

Accepted Manuscript

Title: New evidence about the subduction of the Copiapó ridge beneath South America, and its connection with the Chilean-Pampean flat slab, tracked by satellite GOCE and EGM2008 models

Author: Orlando Alvarez Mario Gimenez Andres Folguera
Silvana Spagnotto Emilce Bustos Walter Baez Carla
Braitenberg



PII: S0264-3707(15)30002-8
DOI: <http://dx.doi.org/doi:10.1016/j.jog.2015.08.002>
Reference: GEOD 1382

To appear in: *Journal of Geodynamics*

Received date: 30-8-2014
Revised date: 11-8-2015
Accepted date: 12-8-2015

Please cite this article as: Alvarez, O., Gimenez, M., Folguera, A., Spagnotto, S., Bustos, E., Baez, W., Braitenberg, C., New evidence about the subduction of the Copiapó ridge beneath South America, and its connection with the Chilean-Pampean flat slab, tracked by satellite GOCE and EGM2008 models, *Journal of Geodynamics* (2015), <http://dx.doi.org/10.1016/j.jog.2015.08.002>

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1 **New evidence about the subduction of the Copiapó ridge beneath South America,**
2 **and its connection with the Chilean-Pampean flat slab, tracked by satellite GOCE**
3 **and EGM2008 models**

4

5 Orlando Alvarez^{1,2} orlando_a_p@yahoo.com.ar, Mario Gimenez^{1,2}
6 gimmario@gmail.com, Andres Folguera^{3,2} andresfolguera2@yahoo.com.ar, Silvana
7 Spagnotto^{4,2} pampa113@gmail.com, Emilce Bustos⁵ emilcebustos@gmail.com, Walter
8 Baez⁵ focobaez@hotmail.com, and Carla Braitenberg⁶ berg@units.it

9

10 ¹*Instituto Geofísico y Sismológico Ing. Volponi, Universidad Nacional de San Juan,*
11 *Ruta 12, Km 17, CP 5407, San Juan, Argentina.*

12 ²*Consejo Nacional de Investigaciones Científicas y Tecnológicas, CONICET,*
13 *Argentina.*

14 ³*IDEAN - Instituto de Estudios Andinos "Don Pablo Groeber". Departamento de Cs.*
15 *Geológicas - FCEN - Pab. II. Universidad de Buenos Aires.*

16 ⁴*Depto de Física - Universidad Nacional de San Luis. San Luis, Argentina.*

17 ⁵*INENCO-CONICET. UNSa. Av. Bolivia 5150. CP 4400 Salta.*

18 ⁶*Dipartimento di Matematica e Geoscienze, Università di Trieste, Via Weiss 1, 34100*
19 *Trieste, Italy.*

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22 **Corresponding Author:** Orlando Alvarez23 **Postal Address:** Mariano Moreno 240(S) dpto:6. Rivadavia, San Juan, Argentina.24 **Phone:** +0054-0264-426737125 **ABSTRACT**

26 Satellite-only gravity measurements and those integrated with terrestrial observations
27 provide global gravity field models of unprecedented precision and spatial resolution,
28 ~~which allow analyzing lithospheric structure~~ allowing the analysis of the lithospheric
29 structure. We used the model EGM2008 (Earth Gravitational Model) to calculate the
30 gravity anomaly and the vertical gravity gradient in the South Central Andes region,
31 correcting these quantities by the topographic effect. Both quantities show a spatial
32 relationship between the projected subduction of the Copiapó aseismic ridge (located at
33 about 27° 30' S), its potential deformational effects in the overriding plate, and the Ojos
34 del Salado-San Buenaventura volcanic lineament. This volcanic lineament constitutes a
35 projection of the volcanic arc towards the retroarc zone, whose origin and development
36 were not clearly understood. The analysis of the gravity anomalies, at the extrapolated
37 zone of the Copiapó ridge beneath the continent, shows a change in the general NNE-
38 trend of the Andean structures to an ENE-direction coincident with the area of the Ojos
39 del Salado-San Buenaventura volcanic lineament. This anomalous pattern over the
40 upper plate is interpreted to be linked with the subduction of the Copiapó ridge.

41 We explore the relation between deformational effects and volcanism at the northern
42 Chilean-Pampean flat slab and the collision of the Copiapó ridge, on the basis of the

2

43 Moho geometry and elastic thicknesses calculated from the new satellite GOCE data.
44 Neotectonic deformations interpreted in previous works associated with volcanic
45 eruptions along the Ojos del Salado-San Buenaventura volcanic lineament is interpreted
46 as caused by crustal doming, imprinted by the subduction of the Copiapó ridge,
47 evidenced by crustal thickening at the sites of ridge inception along the trench.

48 Finally, we propose that the Copiapó ridge could have controlled the northern edge of
49 the Chilean-Pampean flat slab, due to higher buoyancy, similarly to the control that the
50 Juan Fernandez ridge exerts in the geometry of the flat slab further south.

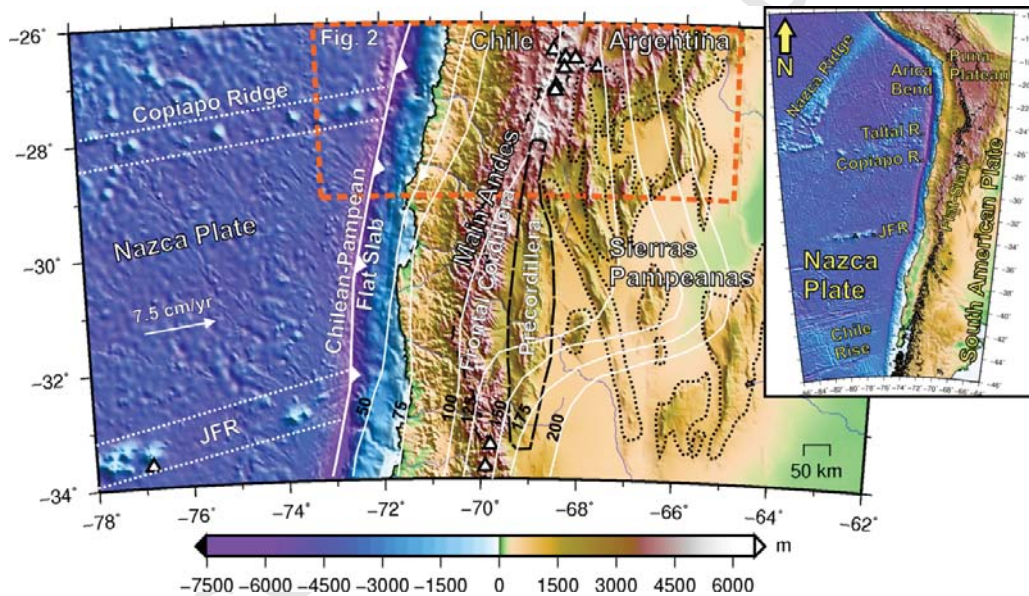
51 **Keywords:** GOCE; Vertical Gravity Gradient; Ojos del Salado-San Buenaventura
52 Lineament; Copiapó aseismic ridge; Chilean-Pampean flat slab.

53

54 1. INTRODUCTION

55 The Andes are the largest active orogenic system developed by subduction of oceanic
56 lithosphere. This continuous and complex mountain belt is the expression of
57 deformational and magmatic processes that acted together with variable intensity along
58 the plate margin. A variable obliquity of the relative convergence between the
59 converging plates, spatio-temporal variable segments of flattenings and steepenings of
60 the subduction zone, delamination phenomena, pluvial gradients and local sites of
61 accretion of oceanic features have determined a strong segmentation of the mountain
62 morphology (Fig. 1) (Gutscher et al. 2000; James and Sacks, 1999; Kay and Coira,
63 2009; Lamb and Davis, 2003; Martinod et al. 2010; Ramos, 2009).

64 Among these controls, shallow to flat subduction settings are of special interest since
 65 they coincide with segments of high mountains and high orogenic amplitude. These
 66 occur at about 10% of the modern convergent margins (Gutscher et al. 2000) and are
 67 commonly associated with subduction of overthickened and therefore abnormally
 68 buoyant oceanic crust, in oceanic plateaus, seamounts, and aseismic ridges. In
 69 particular, the last are considered the major factor associated with the development of
 70 shallow subduction configurations in the Andes (Cloos, 1992; Scholtz and Small, 1997;
 71 Yañez et al. 2001; Yañez and Cembrano, 2004).



72
 73 **Figure 1.** Shaded digital elevation model of the Southern Central Andes region
 74 with superimposed contours of the Wadati-Benioff zone at 25 km intervals (white
 75 solid line) describing the Chilean-Pampean flat slab region (Mulcahy et al.
 76 2014). The Precordillera, depicted by a (black dashed line) and the Sierras
 77 Pampeanas broken foreland, by a (black dotted line), are mountain systems
 78 associated with the development of the flat slab in the last 17 My. Triangles
 79 indicate the current position of the active volcanic arc beyond the limits of the

80 *flat slab (Siebert and Simkin, 2002). An arc gap is associated with the area of*
81 *shallowing of the subducted Nazca plate at these latitudes. The Juan Fernandez*
82 *(JFR) and the Copiapó ridges are indicated (white dotted lines) colliding*
83 *against the Chilean trench at the southern and northern edges of the flat slab*
84 *respectively. Nazca–South American plates convergence rate and azimuth (white*
85 *arrow) is from DeMets et al. (2010).*

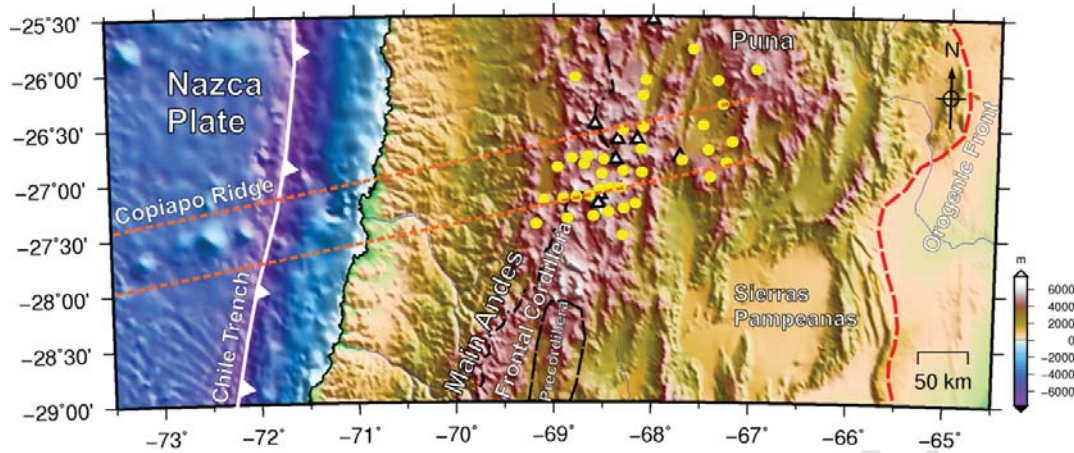
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87 ~~Modern magnitudes of shallow subduction zones are strikingly variable worldwide,~~
88 ~~being the largest and with greater deformational effects located in the Andes (James and~~
89 ~~Sacks, 1999; Kay and Coira, 2009; Ramos and Folguera, 2009). Barazangi and Isacks~~
90 ~~(1976, 1979) related two of the four arc gaps along the Andes to flat slab subduction~~
91 ~~settings based on location of earthquake's hypocenters. In their analysis, the Peru (3°S~~
92 ~~to 15°S) and Chilean-Pampean flat slabs (27°S to 33°S) are tracked beneath the~~
93 ~~continent for more than 700 km from the trench to the foreland zone (McGeary et al.~~
94 ~~1985; Sacks, 1983). In particular, † The Chilean-Pampean flat slab has been linked to a~~
95 ~~broken foreland zone and an eastward expansion of the Neogene to Quaternary arc~~
96 ~~volcanism that produced an arc gap between 28°S to 33°S (Fig. 1) (Allmendinger et al.~~
97 ~~1990; Barazangi and Isacks, 1976, 1979; Cahill and Isacks, 1992; Jordan et al. 1983a,~~
98 ~~1983b; Kay et al. 1988, 1991; Kay and Abbruzzi, 1996; Pilger, 1981; Ramos et al. 1991,~~
99 ~~2002; Smalley and Isacks, 1987; Stauder, 1973; Yañez et al. 2001).~~

100 ~~Crustal seismicity reflects mountain development over flat subduction settings along the~~
101 ~~Central Andes (Kay and Coira, 2009; Ramos et al. 2002). In particular, Pardo et al.~~
102 ~~(2002) refined the shape of the downgoing Nazca Plate from local and teleseismic~~

103 events beneath the Chilean-Pampean flat slab, showing that at the point of inception of
104 the Juan Fernandez ridge (*JFR*), the plate penetrates to a depth of approximately 100-
105 120 km, with a dip of 30°E , flattening underneath the base of the continental lithosphere
106 for several hundreds of kilometers in coincidence with the Precordillera-Sierras
107 Pampeanas zone (26° - 33°S) (Fig. 1). This slab geometry caused the volcanic arc to
108 migrate as far as 600 km away from the trench (Booker et al. 2004; Martinod et al.
109 2010). The Juan Fernández ridge at $\sim 32.5^{\circ}\text{S}$ coincides with a strong bend of the
110 subducted slab depth contours, at the site of truncation of the active volcanic arc in the
111 models of Barazangi and Isacks (1976, 1979) and Pardo et al. (2002). More recently,
112 this geometry of the southern subducted slab has been refined, showing a more
113 symmetrical geometry around the prolonged eastern edge of the subducted ridge
114 (Anderson et al. 2007), showing the direct relation between the slab flattening and the
115 ridge subduction (Fig. 1).

116 At the northern edge of the Chilean-Pampean flat slab, the Copiapó ridge (Contreras-
117 Reyes and Carrizo, 2011), also known as Easter line (Bonatti et al. 1977) collides
118 with the Chilean trench (Fig. 1 and 2). Even though the potential relation between this
119 feature and the northern shallowing of the Nazca plate at these latitudes has been
120 loosely evaluated in different works (Baker et al. 1987; Bonatti et al. 1977; Comte et
121 al. 2002; González-Ferrán et al. 1985), no geophysical data have been used to
122 determine the track of the Copiapó ridge beneath the South American Plate. We explore
123 in this work the potential role of the Copiapó ridge in the described deformational and
124 volcanic patterns of the overriding plate and the definition of the northern flat slab,
125 through its tracking beneath the continent using gravity data.



126

127 **Figure 2.** Shaded digital elevation model at the northern part of de Chilean-Pampean
 128 flat slab (see figure 1 for location). White triangles show the southern extent of the
 129 active volcanic arc at the transition zone between the 30°E dipping Nazca plate and the
 130 flat slab to the south where a gap in arc activity is established. Red dashed line
 131 indicates the eastern orogenic front. Yellow circles depict individual centers of the Ojos
 132 del Salado-San Buenaventura volcanic lineament that develop from the arc zone to the
 133 retroarc area (Kay et al. 2008; Mpodozis et al. 1996; Seggiaro et al. 2000). Note the
 134 general match between the extrapolation of the Copiapó ridge beneath the South
 135 American Plate and the Ojos del Salado-San Buenaventura volcanic lineament.

136

137 We calculated the gravity anomaly and vertical gravity gradient for the South Central
 138 Andes and adjacent offshore region using the relative high spatial resolution of the
 139 global gravity field model EGM2008 (Earth Gravitational Model) and the homogeneous
 140 precision of GOCE (Gravity Field and Steady-State Ocean Circulation Explorer)
 141 satellite data. Moreover, from the Bouguer anomaly calculated from GOCE data, we
 142 computed the crust-mantle discontinuity and the elastic thickness (in the frame of the

7

143 isostatic lithospheric flexure model applying the method of the convolution approach),
144 with the aim of comparing how these parameters vary in both ridge collisions at the
145 northern and southern ends of the flat slab respectively.

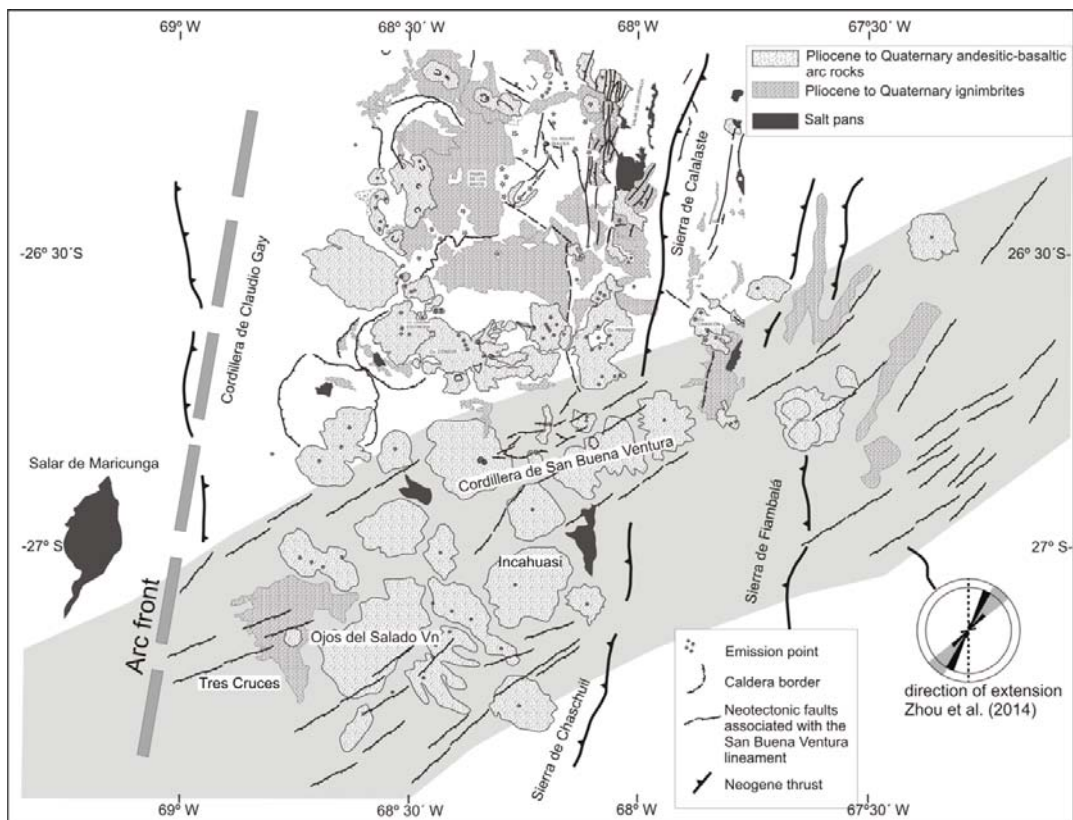
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147 **2. GEOLOGICAL SETTING OF THE NORTHERN CHILEAN-PAMPEAN** 148 **FLAT SLAB (26-28°S)**

149 At the latitudes of the Chilean-Pampean flat slab (approx. 28°S to 32°S) the Andes are
150 formed by a series of mountain systems (Figs. 1 and 2). Namely, from west to east: the
151 Coastal Cordillera, the Frontal Cordillera (which represents the main Andes
152 deformational front), the Precordillera and the Sierras Pampeanas in the foreland. ~~The~~
153 ~~main Andes project to the north beyond the Chilean Pampean flat slab into the southern~~
154 ~~Altiplano region developed at the inner part of the northern Argentinean-Chilean,~~
155 ~~Bolivia and southern Perú fold and thrust belt (Fig. 2). The main Andes are produced by~~
156 ~~mechanisms that involve inversion of Late Triassic extensional detachments and~~
157 ~~subordinately decollement in the Mesozoic and Cenozoic sections (Ramos et al. 1996).~~
158 ~~The age of these uplifts has been determined as early as late Miocene based on the age~~
159 ~~of associated synorogenic strata (Jordan and Allmendinger 1986).~~ The Frontal
160 Cordillera and the relief exposed to the east have been related to the inception of the
161 Chilean-Pampean flat slab between 27-33°S (Ramos et al. 2002). In particular, the
162 Precordillera within this sector is an imbricate system that reactivated Late Triassic
163 detachments and decollements in the Early Paleozoic series, and that together with the
164 Sierras Pampeanas to the east, have been incorporated into the orogenic wedge in the
165 last 10 My (see Ramos et al. 2002, for further references). The Sierras Pampeanas in

166 particular correspond to a classical broken foreland where metamorphosed Paleozoic
167 rocks are exhumed in a Laramide-like way.

168 These structures have coexisted with an oblique convergence regime between the
169 subducted Nazca and South American plates (Fig. 1) ($\sim 77^\circ\text{N}$ respect to the trench)
170 (Angermann et al. 1999; DeMets et al. 1990, 1994, 2010; Kendrick et al. 2003; Ranero
171 et al. 2006; Vigny et al. 2009; Völker et al. 2006). This oblique convergence setting has
172 led to propose a strain partitioned regime favored by the high coupling produced by the
173 flat slab (Chemeda et al. 2000; Gutscher, 1999b, 2000; Pinet and Cobbold, 1992;
174 Pubellier and Cobbold, 1996). In particular, oblique to parallel-to-the-trench
175 regional structures presently act as crustal-scale lateral ramps, partitioning the
176 deformation (Aubry et al. 1996; Baldis and Vaca, 1985; Bassi, 1988; Rossello et
177 al. 1996; Salfity, 1985; Segerstrom and Turner, 1972; Urreiztieta, 1996;
178 Urreiztieta et al. 1996,). Over the Chilean-Pampean flat slab segment, the NNE
179 Calama-Olacapato-El Toro, the Catamarca and the Tucumán lineaments have been
180 associated with strike-slip displacements segmenting the fold and thrust belt (Baldis et
181 al. 1976; Mon, 1976). These structures have also been related to magmatic paths for
182 arc and retroarc volcanism, producing long volcanic alignments in Miocene to
183 Quaternary times (see Kay et al. 2008 and Kay and Coira, 2009 for recent revisions).



184

185 **Figure 3.** Pliocene to Quaternary volcanism and related neotectonic structures
 186 at the Ojos del Salado-San Buenaventura volcanic lineament shaded in grey
 187 (based on Marret et al. 1994; Mpodozis et al. 1996; Seggiaro et al. 2000; Zhou
 188 et al. 2014).

189

190 In particular, the ENE-trending Ojos del Salado-San Buenaventura volcanic
 191 lineament (Zentilli, 1974) extends at $\sim 27^{\circ}\text{S}$ for almost 300 km from the arc front
 192 to the retroarc zone (Figs. 2 and 3) (Carter, 1974; Gerth, 1955). This lineament
 193 starts at the Chilean Central Valley and crosses the Andes producing a major
 194 transverse morphological discontinuity composed of the highest volcanoes of the
 195 world such as the Ojos del Salado dome complex and Tres Cruces and Incahuasi

10

196 stratovolcanoes (Fig. 3) (Bonatti et al. 1977; Mpodozis and Kay, 1992;
197 Mpodozis et al. 1996). This volcanic chain is associated with a strong deflection
198 in the Andean drainage divide from N- in the south to ENE in the north imposed
199 by the aligned summit of stratovolcanoes, calderas and dome complexes (Fig. 2)
200 (Kay et al. 2008). (see Appendix A for a more detailed description) ~~These centers are~~
201 ~~affected and controlled by normal and right lateral neotectonic structures,~~
202 ~~defining a general ENE structural alignment (Fig. 3) (Kay et al. 2008; Marrett et~~
203 ~~al. 1994; Mpodozis et al. 1996; Seggiaro et al. 2000; Zhou et al. 2014). Most of~~
204 ~~the volcanic centers controlled by these structures are Pleistocene in age,~~
205 ~~although some show a more prolonged activity during the Holocene (see~~
206 ~~Appendix A for a more detailed description) (Baker et al. 1987; Kraemer et al. 1999;~~
207 ~~Montero López, 2009; Mpodozis et al. 1996; Risse et al. 2008; Seggiaro et al.~~
208 ~~2000; Viramonte et al. 2008).~~ → moved to Appendix A

209

210 **3. METHODOLOGY: EARTH GRAVITY FIELD MODELS**

211 Satellite only models (e.g. GOCE, Pail et al. 2011) and high-resolution gravity field
212 models based on observations of satellite data plus terrestrial data are available in
213 spherical harmonic expansion (e.g., Earth Gravitational Model 2008, EGM2008 Pavlis
214 et al. 2008). ~~The resolution of the Satellite-only models present is lower resolution than~~
215 ~~terrestrial data and combined models, although it is they are accurate for interpretation~~
216 of large-scale structures, especially in high and inaccessible regions such as certain parts
217 of the Andes (Köther et al. 2012). This allows regional gravity modeling, for studies of
218 the lithosphere and crustal anomalies such as regional structures, suture zones, basins

219 and delimitation of magmatic provinces (Álvarez et al. 2012, 2014a,b, 2015;
220 Braitenberg et al. 2011a, 2015; Hirt et al. 2012; Li et al. 2013; Mariani et al. 2013)
221 especially in those regions where terrestrial data are sparse or unavailable due to the
222 limited access. The use of satellite derived models also allows overcoming problems
223 related to non-unified height measurements from different terrestrial campaigns (Gatti et
224 al. 2013; Reguzzoni and Sampietro, 2010). Earth Gravity Field models are presented as
225 sets of coefficients of a spherical harmonic approximation of the gravity field up to a
226 maximum degree N_{\max} which governs the spatial resolution of the model (λ)
227 (Barthelmes, 2009; Hofmann-Wellenhof and Moritz, 2006; Li, 2001). EGM2008 is
228 developed up to degree/order 2159 with some additional terms up to degree/order 2190
229 (Pavlis et al. 2008). The last model derived from data of the satellite GOCE mission
230 (datasheet_go_cons_gcf_2_tim_r4.pdf, <http://icgem.gfz-potsdam.de/ICGEM/>, Pail et al.
231 2011) is developed up to a maximum degree and order $N=250$ (with an effective data
232 volume of approx. 26.5 months). Thus, the spatial resolution for the potential field
233 model EGM2008 becomes equal to $\lambda/2 \approx 9$ km, and $\lambda/2 \approx 80$ km for GOCE.

234 The gravity field is attenuated at the satellite height, so satellite-only based models
235 achieve a low spatial resolution and provide information only on the long wavelength
236 portion of the spectrum (Reguzzoni and Sampietro, 2010). Despite this disadvantage,
237 GOCE derived models have homogeneous precision, as these have no present sampling
238 errors and biases induced by terrestrial data, as in the EGM2008 model. Statistical
239 results from a comparison analysis between gravity anomalies obtained from both
240 models (Appendix B), up to the same degree/order of the harmonic expansion ($N=250$),
241 indicate that the fields are only in partial agreement and differences become smooth
242 (Braitenberg et al. 2011a; Alvarez et al. 2012, 2013).

243 In particular, gravity anomaly and vertical gravity gradient maps obtained from satellite
244 GOCE data in the northern Chilean-Pampean flat slab exhibit a good correlation at
245 medium to high wavelengths with the EGM2008 model, especially over the areas of
246 low topography (Alvarez et al. 2012, 2013). Greater differences between the EGM2008
247 with respect to the satellite-only model of GOCE arise in the Andean region. North of
248 28°S, the high gravimetric effect of the Andean roots masks the crustal structures when
249 using the long wavelength data of GOCE (Alvarez et al. 2012).

250 Since this work is focused on the evaluation of the interplay between the subduction of
251 an aseismic ridge (Copiapó ridge) and the potential deformational effects in the
252 overriding plate and volcanism along the Ojos del Salado-San Buenaventura volcanic
253 lineament, a gravity data set with a regional coverage without neglecting high frequency
254 data is required. EGM2008 model constitutes a combined solution composed of a
255 worldwide surface gravity anomaly database of 5' x 5' resolution, and GRACE-derived
256 satellite solutions, that has the advantage of covering both land and marine areas
257 worldwide. Pavlis et al. (2012) indicate that its spectral content was supplemented with
258 gravitational information implied by the topography, only over areas where lower
259 resolution gravity data were available. Over these “fill-in” areas, the gravity anomaly
260 information over the harmonics of degree from 721 to 2159 is supplemented with the
261 gravitational information obtained from the analysis of a global set of Residual Terrain
262 Model-Implied (RTM-Implied) gravity anomalies (i.e. the high frequencies of
263 EGM2008 are calculated from the topography; see Pavlis et al. (2012), for a detailed
264 discussion). The quality and resolution of the downward continued geopotential models
265 in the Andes and Central America decrease with increasing topography and depend on
266 the availability of terrestrial gravity data (Köther et al. 2012). In the study area the 5' x

267 5' gravity anomaly (Δg) data sources are NGA-LSC (National Geospatial-Intelligence
268 Agency - Least-Squares Collocation) and some fill-in data in areas covered by
269 proprietary data. The 5'x 5' Δg data availability shows several lines of available data in
270 our study region (Pavlis et al. 2008). Köther et al. (2012) explained that combined
271 gravity models can be used for density modeling of relatively smaller features such as
272 shallower crustal structures, while satellite-only models are not appropriate for this
273 purpose due to their low spatial resolution. They also mentioned that 3D modeling of
274 synthetic gradients and invariants of subduction zones, using the Andes as case study,
275 proved the applicability of gradient measurements for detection of the edge of
276 geological structures. Therefore, gradients from GOCE mission can resolve structural
277 information (Köther et al. 2012).

278 Based on the aforementioned, the T_{zz} map of EGM2008 data is preferred to GOCE in
279 order to identify the main structures that affect the upper crust, having a higher spatial
280 resolution. However, in order to calculate different quantities related to a longer
281 wavelength characteristic of the gravity field (Moho and Elastic Thickness) we used the
282 satellite-only model of GOCE which presents homogeneous precision and higher
283 accuracy (see Alvarez et al. 2012, 2014, 2015).

284

285 **3.1. Gravity anomaly and vertical gravity gradient fields for identifying crustal** 286 **regional structures**

287 The topography-reduced gravity anomaly (Ga eq. 1) (Hofmann-Wellenhof and Moritz,
288 2006; Molodensky et al. 1945, 1962), a functional of the geopotential (not defined by
289 reduction formulas such as the Bouguer anomaly), is very useful to show the effects of

290 different rock densities within the crust. It is calculated by subtracting the gravity of the
 291 reference potential from the observed gravity and then subtracting the effect of the
 292 topographic masses above the geoid (see Barthelmes, 2009).

$$293 \quad Ga = \Delta_{g_T}(\mathbf{h}, \lambda, \phi) = g(\mathbf{h}, \lambda, \phi) - g_r(\mathbf{h}, \lambda, \phi) - \gamma(\mathbf{h} - \zeta_g, \phi) \quad (1)$$

294 Where g is the real gravity, g_r is the effect of the topographical masses above the geoid at
 295 a given point (h, λ, ϕ) , and γ is the gravity of the reference potential at the same
 296 longitude and latitude, but at the ellipsoidal height $h - \zeta_g$ (where ζ_g is the generalized
 297 height anomaly). The height h is assumed on or outside the Earth surface ($h \geq h_i$).

298 Gradients of the gravity field highlight main geological features, and allow unveiling
 299 unknown structures that are either buried by sediments or just have not been recognized
 300 (Braitenberg et al. 2011a). The Marussi Tensor M is composed of five independent
 301 elements and is obtained as the second order spatial derivatives of the disturbing
 302 potential (Hofmann-Wellenhof and Moritz, 2006). The Marussi tensor components in a
 303 spherical coordinate system are given by Tscherning (1976) and by Rummel et al.
 304 (2011).

305 For geological mapping, the second vertical derivative of the disturbing potential (T_{zz}
 306 component eq. 2) is ideal, since it highlights the center of the anomalous mass
 307 (Braitenberg et al. 2011a). Braitenberg et al. (2011a) explained that even though the
 308 Marussi tensor and the gravity field both reflect density variations in the crust, these
 309 outline very different subsurface features. In a recent work, Alvarez et al. (2012)
 310 showed that both quantities highlight equivalent geological features in a different and
 311 complementary way: while T_{zz} highlights mass heterogeneities, where density contrasts
 312 are faced, especially in the upper crust, the Ga becomes useful when the density contrast

313 is relatively low and the geological structures are deep, where the T_{zz} loses sensitivity
 314 (Alvarez et al. 2012).

$$315 \quad T_{zz} = \frac{\partial^2 T}{\partial r^2} \left[1 \text{ Eötvös} = 10^{-4} \frac{\text{mGal}}{\text{m}} \right] \quad (2)$$

316

317 **3.1.1. Calculation of G_a and T_{zz}**

318 We calculated the vertical gravity gradient and the gravity anomaly (Figs. 4a and b)
 319 (Janak and Sprlak, 2006) for the Southern Central Andes, using data of the model
 320 EGM2008 up to degree/order $N=2159$ (Pavlis et al. 2008) on a regular grid with a cell
 321 size of 0.05° . The values were calculated in a geocentric spherical coordinate system at
 322 a calculation height of 7,000 m to ensure that all values were above the topography.

323 The topographic effect was removed from the fields to eliminate the correlation with the
 324 topography. Topographic mass elements obtained from a global relief model, which
 325 includes ocean bathymetry (ETOPO1: Amante and Eakins, 2008), were approximated
 326 with spherical prisms of constant density in a spherical coordinate system (Anderson,
 327 1976; Grombein et al. 2010, 2013; Heck and Seitz, 2007; Wild-Pfeiffer, 2008).
 328 Spherical prisms are used in order to account the Earth's curvature (Alvarez et al. 2013;
 329 Bouman et al. 2013; Uieda et al. 2010). A standard density of $2,670 \text{ kg/m}^3$ was used for
 330 masses above sea level and a density of $1,030 \text{ kg/m}^3$ for the sea water. All calculations
 331 were carried out with respect to the system WGS84. The topographic correction
 332 amounts up to tens of Eötvös for the vertical gradient and up to a few hundreds of mGal
 333 for gravity, becoming higher over the maximum topographic elevations (e.g. the Puna

334 and the Main Andes) and lower over the topographic depressions such as the Chilean
335 trench.

336

337 **3.2. Flexural strength of the lithosphere**

338 The lithospheric strength, related to the thermal state, composition and rheological
339 properties of the crust is well characterized by the flexural rigidity (Lowry et al. 2000).

340 The flexural rigidity (D) can be interpreted in terms of the elastic thickness (Te) by
341 making assumptions regarding the Poisson ratio ($\nu=0.25$) and the Young modulus
342 ($E=100\text{GPa}=10^{11}\text{N/m}^2$). Thus, the flexural rigidity can be expressed by:

$$343 \quad D = T_e^3 \cdot \frac{E}{12(1 - \nu^2)} \quad (3)$$

344 The isostatic state and deformation of the upper crust are reflected in the spatial
345 distribution of the Te , whose variation can be explained by temperature distribution and
346 a change in the Young modulus. Te defines the maximum size and wavelength of the
347 surface loading that can be supported without an elastic break of the lithosphere.

348 Different authors found a correlation between Te and the geometry and composition of
349 the flexured plate, external forces and the thermal structure (e.g. Burov and Diament,
350 1995; Goetze and Evans, 1979; Hackney et al. 2006; Lyon-Caen and Molnar, 1983).

351 Different methods have been developed, tested and extensively used for estimation of
352 the elastic thickness, as flexural coherence analysis or spectral methods, and recently the
353 convolution approach (Braitenberg et al. 2002, Wienecke 2006, 2007). When
354 calculating the flexural strength of the lithosphere using spectral methods (coherence

355 and admittance) a large spatial window is required over the study area, and the method
356 becomes unstable if the input topography is smooth. Both methods require an averaging
357 process; therefore the variation in rigidity may be retrieved only to a limited extent
358 (Wienecke 2006). For this reason, these techniques have been questioned when applied
359 to continental lithosphere.

360 The convolution approach is a method of double entry that calculates the flexure
361 parameters by the best fit of the observed crust-mantle interface (e.g. Moho by gravity
362 inversion) and the crust-mantle interface computed due to a flexure model
363 (www.lithoflex.org). This method requires gravity field data on a much smaller scale
364 (on the order of 100 km in length) than when using spectral methods, and allows a
365 relatively high spatial resolution (Braitenberg et al. 2007; Wienecke, 2006). Otherwise,
366 the topography must be known over an extensive scale, which depends on the elastic
367 thickness and therefore the radius of convolution as explained by Wienecke (2006).
368 This method has been extensively tested in synthetic models over different areas
369 worldwide (Alvarez et al. 2014b, 2015; Braitenberg and Drigo, 1997; Braitenberg and
370 Zadro, 1999; Braitenberg et al. 1997; Braitenberg et al. 2002; Bratfisch et al. 2010;
371 Ebbing et al. 2007; Ferraccioli et al. 2011; Steffen et al. 2011; Wienecke, 2002, 2006;
372 Zadro and Braitenberg, 1997). We used Lithoflex software package to accomplish the
373 inverse modeling of T_e (see Appendix C).

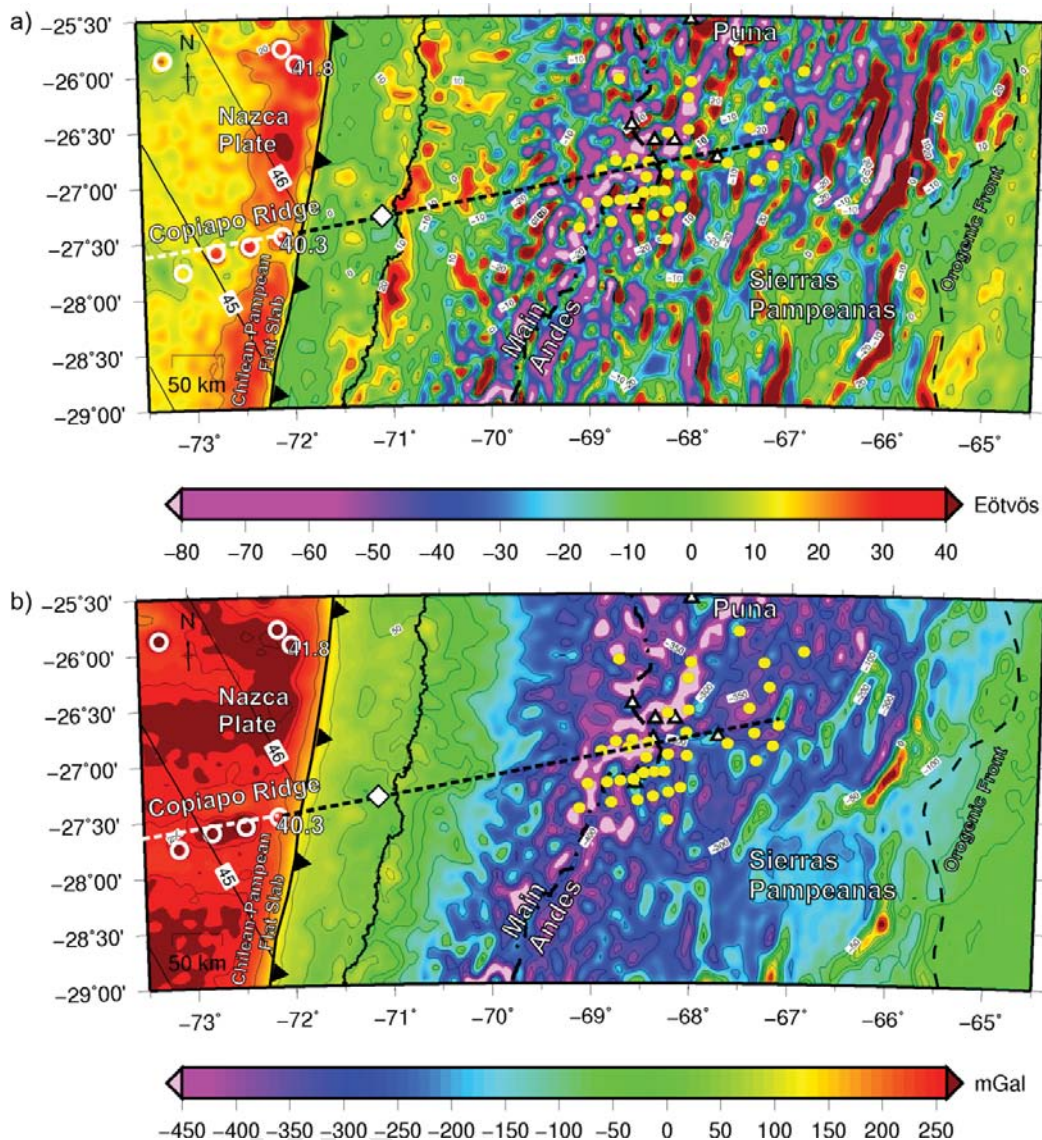
374

375 **4. RESULTS**

376 **4.1. The Copiapó ridge subduction and its deformational effects from EGM2008** 377 **derived data**

378 The outer rise east of the Chilean trench produced by the flexure of the downgoing
379 Nazca plate, coincides with a positive Ga of about 230 mGal (Fig. 4b), and with a
380 positive T_{zz} anomaly higher than 23 Eötvös (Fig. 4a). Next to it, the Copiapó ridge
381 appears as a well-defined T_{zz} signal higher than 30 Eötvös and a Ga higher than 250
382 mGal. We calculated the path of the Copiapó ridge beneath South America using Euler
383 Poles from Table A1 (Wessel and Kroenke, 1997, 1998; Wessel 1999, 2001, 2008. See
384 table A1 in appendix section) determining its projection in 2 My and 8 My onwards
385 from the point of inception into the trench (Fig. 5a). The backward calculation
386 indicates a probable origin at the Easter (Sala and Gomez) hotspot approximately 40
387 My ago.

388 At the latitudes between 26.5° and 27.5°S, over the South American continent, a rotation
389 in the general strike of the NNE pattern of the Ga and T_{zz} anomalies to a general ENE-
390 direction is particularly notorious (Fig. 4a, b). This is interpreted as reflecting a regional
391 rotation in the Andean structures circumscribed to a band coincident with the
392 extrapolation of the Copiapó ridge beneath the South American plate, determined using
393 Euler poles between Nazca and South American plates (Fig. 4 and 5a).



394

395 **Figure 4. a)** Vertical Gravity Gradient from EGM2008 corrected by topographic effect
 396 in the region of the Copiapó ridge and Ojos del Salado-San Buenaventura volcanic
 397 lineament. **b)** Gravity anomaly computed from the EGM2008 corrected by topographic
 398 effect. Superimposed, the computed track of the Copiapó ridge (Fig. 5a) is indicated
 399 backward in time (white dashed line) and forward in time (black dashed line). Yellow
 400 circles depict individual centers forming the Ojos del Salado-San Buenaventura
 401 volcanic lineament (Fig. 2). White circles are seamount locations and age of the

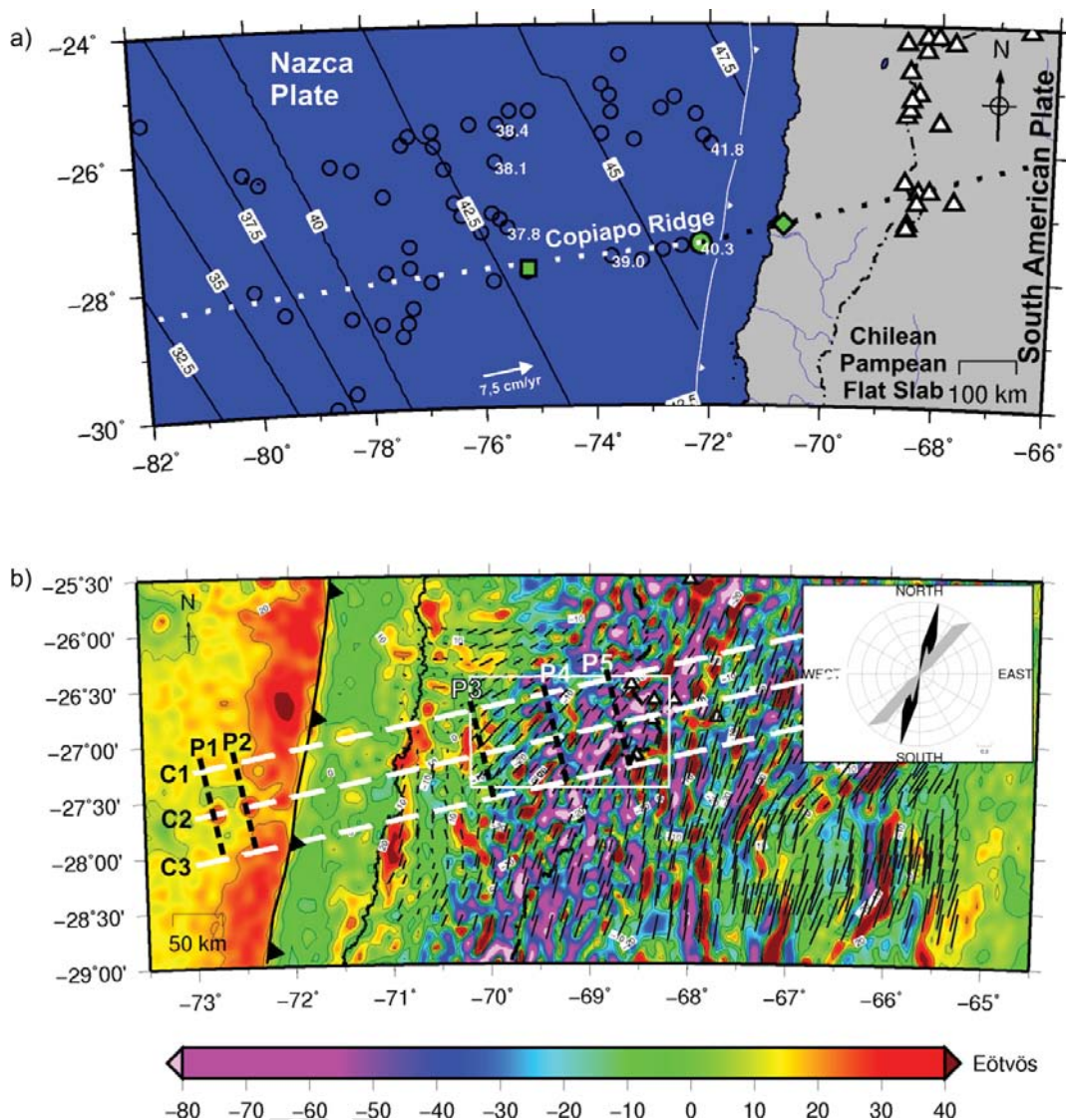
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402 *underlying sea floor in millions of years, from the catalog of Wessel (2001). Solid line*
403 *indicates oceanic crust ages (Müller et al. 2008). Triangles indicate the current*
404 *position of the active volcanic arc (Siebert and Simkin, 2002).*

405

406 The relation between the extrapolation of the Copiapó ridge beneath the South
407 American plate and the change in the general NNE-trend of the Andean structures to an
408 ENE-direction, coincident with the area of the Ojos del Salado-San Buenaventura
409 volcanic lineament, was additionally analyzed by means of the terrain fabric analysis
410 (see Appendix D). Previous works used the eigenvector analysis for detection of
411 lineaments by means of satellite images and digital elevation data (e.g. Koike et al.
412 1998; Raghavan et al. 1993). In this sense, in a recent work, Beiki and Pedersen (2010)
413 used the eigenvector analysis of gravity gradient tensor to locate geologic bodies.

414 We used Microdem software to extract the fabric from the vertical gravity gradient, in
415 order to recognize the structural pattern at the site of subduction of the Copiapó
416 ridge (Fig. 5b). We obtained a circular histogram (rose diagram) from the fabric of the
417 T_{zz} for the whole area (black azimuth line in the rose diagram of Fig. 5b) and another
418 for the region (white rectangle in Fig. 5b) where a general deflection of the Andean
419 structures is observed. From the rose diagram, the deflected fabric (to an ENE-direction)
420 along this stripe (grey azimuth line in Fig. 5b) is interpreted as related to a
421 deformational fabric imposed by the subduction of the Copiapó ridge beneath South
422 America.



423

424 **Figure 5. a)** Reverse and forward reconstruction of the Copiapó ridge trajectories
 425 using Müller et al. (2008)'s and DeMets et al. (2010)'s Euler Poles respectively. The
 426 green squares indicate the corresponding positions 2 My ago for ~~two~~ a seamounts
 427 (green circle) located near the trench, while the green rhombs—indicate ~~their~~ it's
 428 projected position beneath the continent in 2 My. Triangles indicate the current
 429 position of the active volcanic arc (Siebert and Simkin, 2002). Blue circles are
 430 seamount locations with the age of the underlying sea floor (My) from the catalog of
 431 Wessel (2001). Solid lines indicate oceanic crust ages (Müller et al. 2008). **b)** Strength

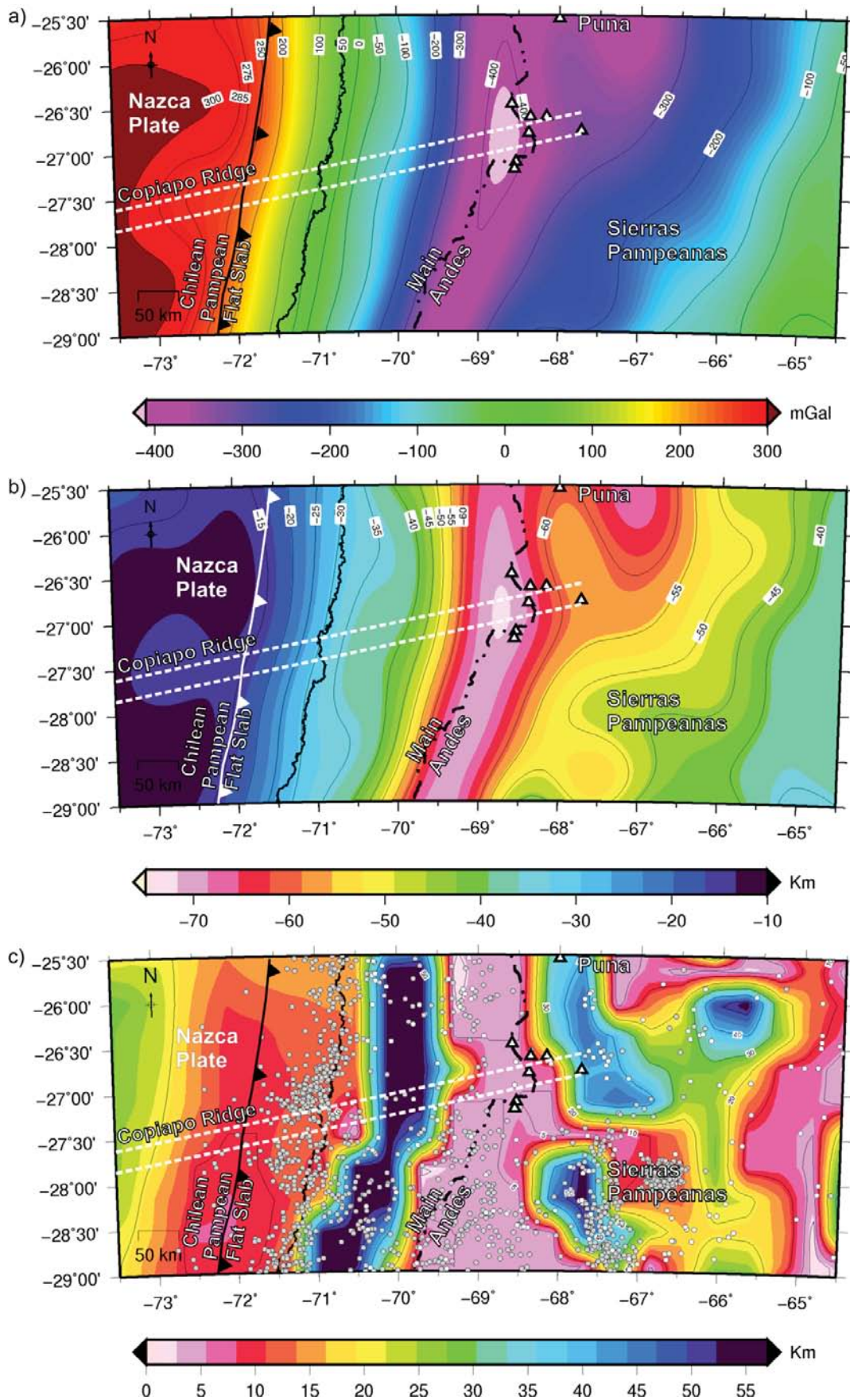
22

432 *and orientation of the fabric of the vertical gradient of the gravity field (T_{zz}),*
433 *superposed to the vertical gradient of the gravity field at the northern part of the*
434 *Chilean-Pampean flat slab zone. Note the strong deflection that is observed at the zone*
435 *of extrapolation of the Copiapó ridge beneath the South American plate. In the upper*
436 *right corner, a rose diagram shows the deviation from the strikes of the vertical*
437 *gradient of the gravity field, for the whole study area (black) to the sector where the*
438 *deflection is inferred to be caused by the Copiapó ridge collision (gray) (white*
439 *rectangle). White dashed lines are profiles (C_i) of Fig. D1. Black dashed lines are*
440 *profiles (P_i) of Fig. D2.*

441

442 **4.2. Strength of the lithosphere from GOCE data**

443 The Bouguer anomaly obtained from GOCE data (Fig. 6a) and the gravity anomaly
444 corrected by the topographic effect from EGM2008 (Fig. 4b) show the influence of the
445 Andean roots expressed by regional low gravity values (less than -200 mGal), being
446 lower in the Puna region to the north of the analyzed segment. The positive effect of the
447 Nazca plate is also observed, reaching maximum values at the outer rise area. The
448 Copiapó ridge can be tracked by its well-defined gravity signal, lower than the
449 surrounding oceanic ~~ocean~~ floor (Fig. 6a). Additionally, Moho depths (Fig. 6b) in the
450 oceanic Nazca Plate show the existence of an over thickened crust in coincidence with
451 the ridge path (~15 km). Similarly, the Moho depths increase beneath the *JFR*, as
452 reported by Von Huene et al. (1997) based on wide angle seismic data and by Sandwell
453 and Smith (1997) who related negative satellite derived gravity anomalies to a crustal
454 root indicative of crustal flexure derived from loading.



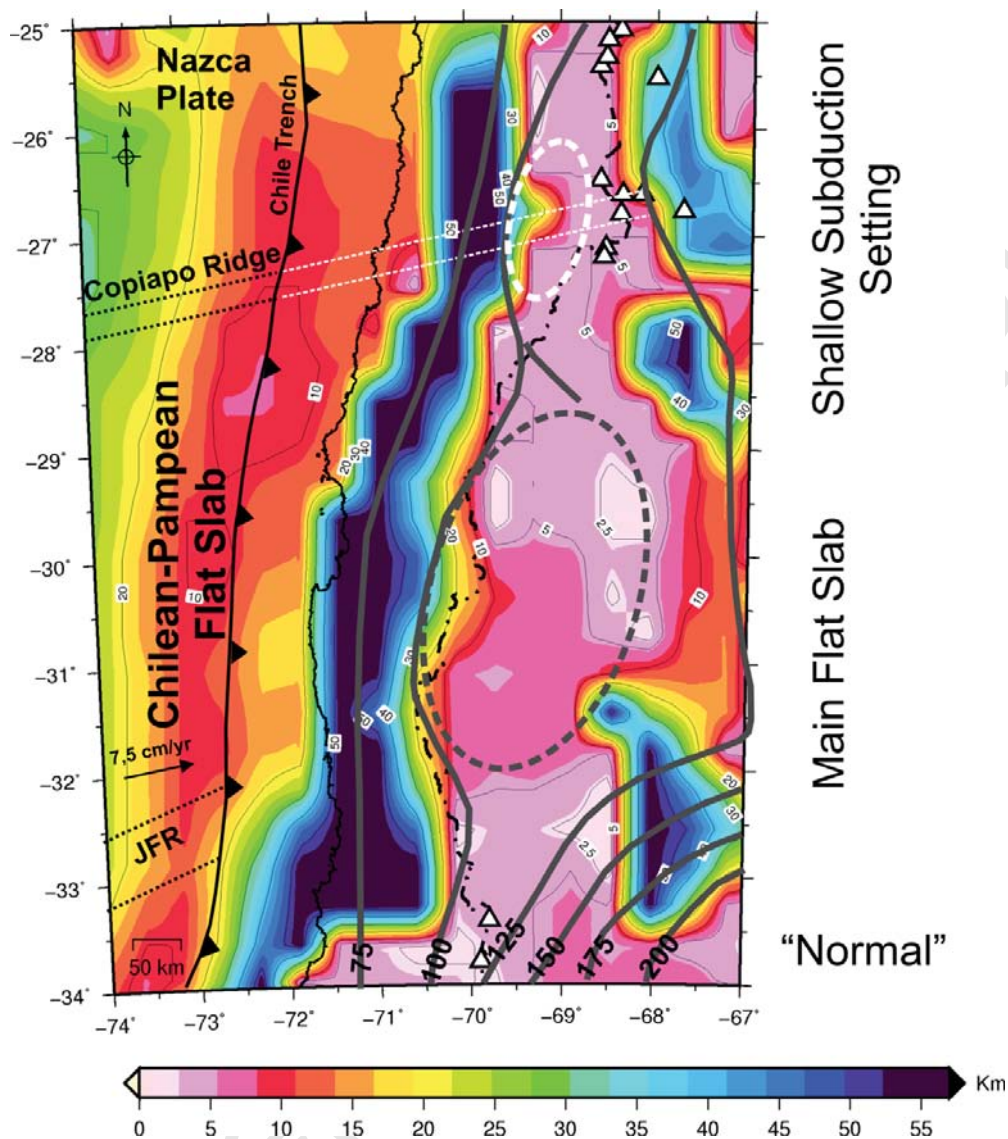
456 **Figure 6. a)** Bouguer anomaly obtained from GOCE (up to degree/order $N=250$)
457 corrected for main sedimentary basins. This anomaly has been used to invert the Moho
458 surface. Note an accentuation of the negative anomaly produced by the Andean roots at
459 the site of inception of the Copiapó ridge beneath South America. The Copiapó ridge
460 over the Nazca plate can be discerned by its well-defined gravity signal, lower than the
461 surrounding ocean floor. **b)** Moho undulations obtained by inversion of the sediment-
462 corrected Bouguer anomaly (GOCE data, Fig. 6a). The obtained Moho depths suggest
463 an over thickened oceanic crust along the ridge path before colliding against the
464 Chilean margin. **c)** Elastic thickness obtained from GOCE data, superimposed to
465 hypocentral seismicity indicated by small circles (EHB Catalog) (crustal earthquakes
466 are in white and subducted Nazca Plate earthquakes in black) compared to the
467 computed path of the Copiapó ridge. Note a seismic gap onshore along the seamount
468 track that coincides with a site of low T_e . Also note how a deepest Moho at the site of
469 interaction between the subducted ridge and the Andean roots is deflected to the east.

470

471 The Bouguer anomaly obtained from GOCE data (Fig. 6a) evidenced a maximum depth
472 of the Andean roots, at the site of inception of the Copiapó ridge beneath South
473 America, reflected in the inverted Moho topography (Fig. 6b). However, ~~a~~ At the
474 forearc region a change in the Moho contours is observed (between -35 to -40 Km)
475 indicating a crust ~~narrowing~~ thickening aligned with the ridge path (Fig. 6b). The same
476 pattern had been reported from seismological data (between -35 to -40 Km) for the *JFR*
477 at the southern part of the Chilean-Pampean flat slab (see Alvarez et al. 2015 and
478 references therein).

479 Crustal rigidity increases in the forearc over the flat-slab region, as noted in previous
480 works (Alvarez et al. 2014b, 2015; Stewart and Watts, 1997; Tassara, 2005) defining a
481 sharp transition to the south of the Juan Fernandez ridge zone of inception. In particular,
482 elastic thicknesses in the oceanic plate next to the trench (Fig. 7) are higher where the
483 Juan Fernandez ridge intersects the trench. South of *JFR* the forearc exhibits low elastic
484 thicknesses interpreted as the expression of a weakened crust due to the heating of the
485 asthenospheric wedge (Alvarez et al. 2015). A correlation between seismicity (EHB-
486 Catalog) and higher T_e values exists along the Juan Fernandez ridge path beneath the
487 South American plate (Alvarez et al. 2015). Also, a close correspondence exists
488 between the contours of the subducted Nazca slab (plate morphology) and T_e
489 variations (higher than 5 km) along the southern region of the flat slab, particularly at
490 the Andes (Alvarez et al. 2014b, 2015).

491 Contrastingly, the Copiapó ridge path beneath South American presents an aseismic
492 behavior (Fig. 6c). In accordance, T_e values suggest a lower rigidity than in the
493 interplate zone along the *JFR* path. Despite this, a strong truncation of the 5 km isoline
494 of T_e along the Copiapó ridge path suggests a higher rigidity respect to its surroundings
495 (Figs. 6c and 7). In order to visualize these changes, we traced three longitudinal cross-
496 sections (see Figure 5b for location) along the path of the Copiapó ridge and to the north
497 and south of it respectively, comparing topography, T_{zz} , Bouguer anomaly and T_e (see
498 Fig. E1 in Appendix E). We also plotted five 2D profiles across the Copiapó ridge path
499 (see Figure 5b for location) comparing: relief, T_{zz} , Bouguer anomaly, Moho and T_e (see
500 Fig. E2 in Appendix E).



501

502 **Figure 7.** Elastic thickness obtained from GOCE data. T_e presents high values at the
 503 forearc in the flat slab region (28.5° to 33°S) with a higher rigidity at the backarc
 504 through the extrapolated path of the JFR. Near the Copiapó ridge, the T_e presents a
 505 different behavior at the forearc, with high T_e values near the trench decreasing
 506 eastwards more rapidly than in the JFR region. In spite of this, along the extrapolated
 507 path of the Copiapó ridge higher T_e values are present, although at a minor scale when
 508 compared to JFR. The Wadati-Benioff zone contours (grey solid line) is represented
 509 after Mulcahy et al. (2014) (dotted grey ellipse approximates the region where the

27

510 *Wadati-Benioff zone is near 100-110 km depth over the Chilean-Pampean flat slab).*
511 *The dotted white ellipse represents the proposal of a shallow subduction segment*
512 *potentially related to the subduction of the Copiapó ridge.*

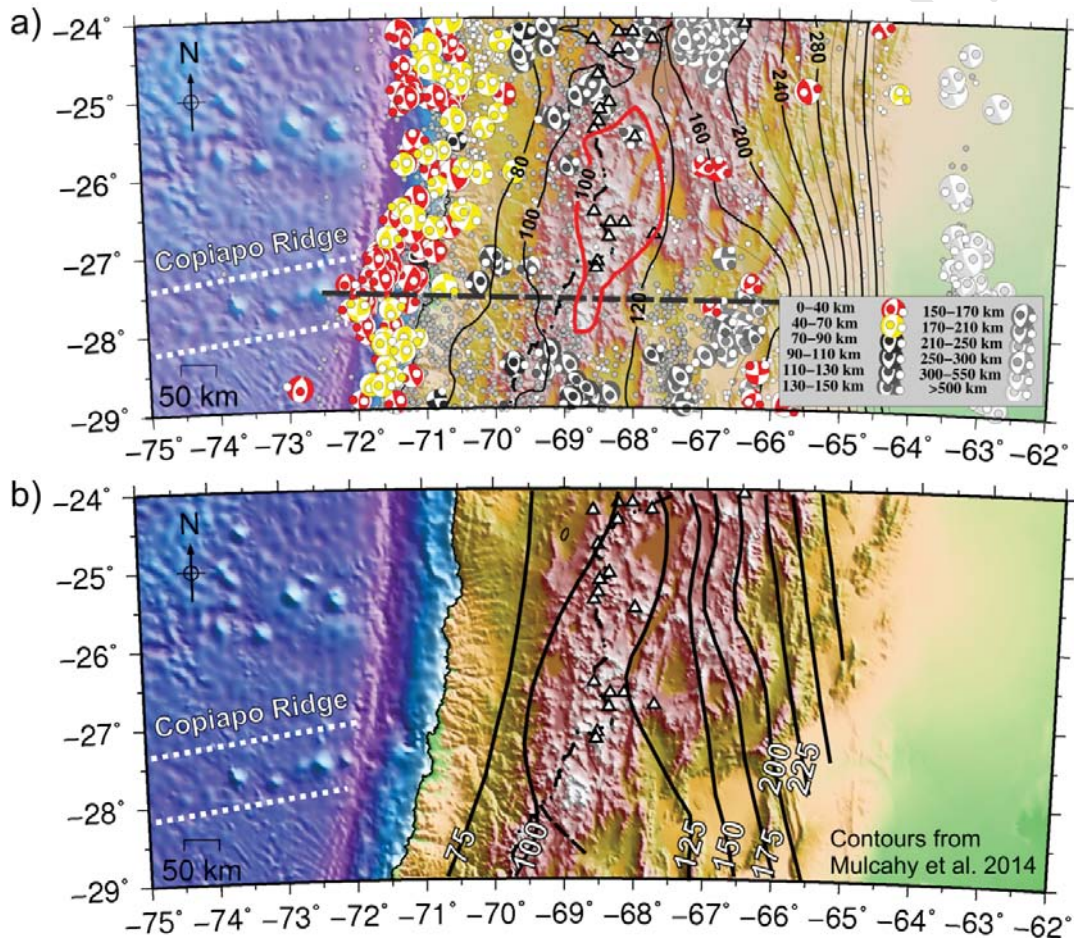
513

514 **4.3. Slab geometry at the site of inception of the Copiapó ridge (26°S to 28°S)**

515 Early works had delineated the northern edge of the Chilean-Pampean flat slab across
516 27-28°S, as a rather smooth feature, from nearly flat in the south to “normal” in the
517 north (~30°E) (Araujo and Suarez, 1994; Bevis and Isacks, 1984; Cahill and Isacks,
518 1992; Jordan et al. 1983a; Pardo et al. 2002; Smalley and Isacks, 1987). More recently
519 new seismic experiments have allowed defining more clearly a sharper transition (Fig.
520 8b) between these segments (Mulcahy et al. 2014).

521 We projected hypocenters within the latitude window of 27.5°S using the CMT
522 locations and EHB-Catalog. Figures 8a and 9a show that the plate subducts sub-
523 horizontally for the first 200 km, and then penetrates into the asthenosphere with an
524 approximate angle of 20°E. The catalog of focal mechanisms determined by the Harvard
525 Centroid Moment Tensor catalog (Harvard CMT, Global CMT Project, 2006) is
526 obtained by means of the centroid moment tensor method (CMT). This consists in the
527 inversion of two parts of the seismogram, 1) body waves of long period and 2) surface
528 waves of very long period (Stein and Wysession, 2003). CMT solutions use complete
529 waveforms, resulting in the centroid, or average location, in space and time, of the
530 released seismic energy, so depth data can be considered reliable. Additionally, the
531 EHB-Catalog (EHB-Catalog, 2009; Engdahl et al. 1998) has a more reliable location of
532 earthquakes in the study region (Fig. 8a).

533 In map view (Fig. 8a), the 160 km iso-depth contour clearly depicts the shallow to flat-
 534 slab geometry getting a more symmetrical shape in map view than in the initial
 535 proposals of Cahill and Isacks (1992). Recent seismic surveys (Mulcahy et al.
 536 2014) using local networks confirm this morphology with a steep edge of the flat
 537 slab near 27°S (Figs. 8b and 9b).

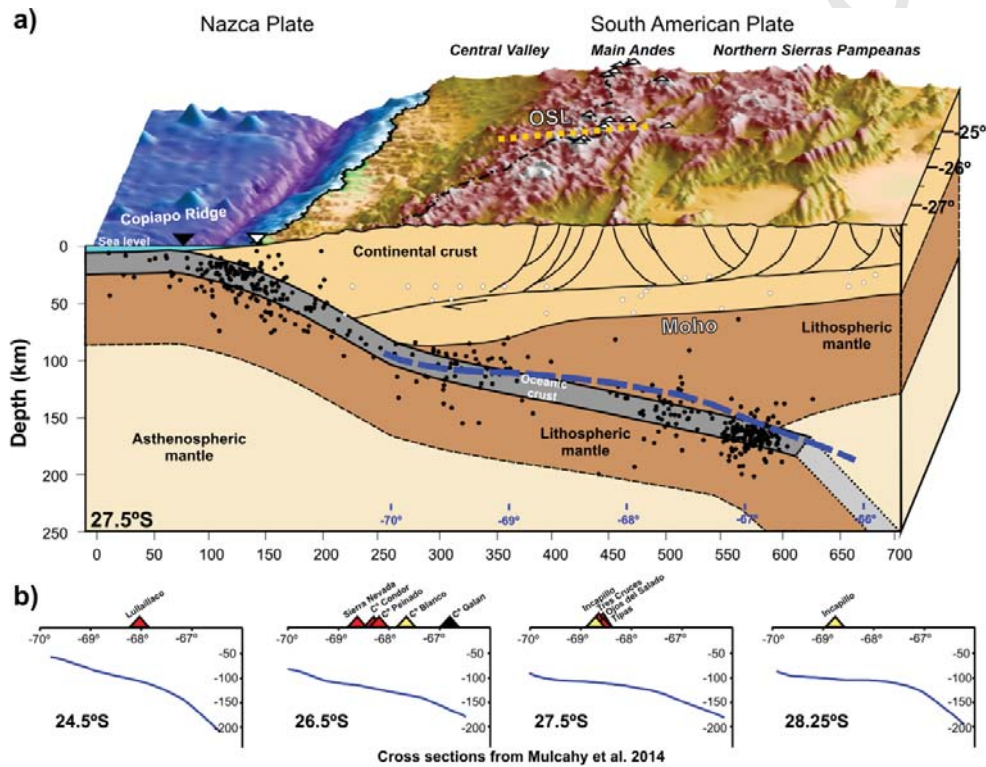


538

539 **Figure 8. a)** Iso-depth contours and focal mechanisms associated with the downgoing
 540 Nazca plate in the northern region of the Chilean-Pampean flat slab and active crustal
 541 structures of the Andes. The 160 km iso-depth contour depicts clearly the Chilean-
 542 Pampean flat-slab geometry. The shallowing of the subducted Nazca Plate observed at
 543 the inception point of the Copiapó ridge is highlighted as a red contour. Hypocentral

29

544 seismicity is indicated with small circles (EHB Catalog) (crustal earthquakes are in
 545 white and subducted Nazca Plate earthquakes in black). Focal Mechanisms are from
 546 CMT Harvard Catalog. Black dashed line indicates cross section in Figure 9. **B)**
 547 Contoured depths of the Wadati-Benioff Zone from Mulachy et al. (2014) at 25 km
 548 intervals, shown in black. Both contour families (a and b) show a shallower subducted
 549 Nazca Plate at the inception point of the Copiapó ridge.



550

551 **Figure 9a.** Digital elevation model in perspective with crustal cross section at 27.5°S
 552 based on Salfity et al. (2005) (azimuth 92.5°; +/- 50 km wide; 0-650 km depth).
 553 Interpreted Wadati-Benioff zone at these latitudes is determined by hypocentral
 554 seismicity indicated with black circles (EHB Catalog). Crustal earthquakes are
 555 indicated as small white circles. As a reference, a cross section at 27.5°S from Mulachy
 556 et al. (2014) is shown (blue dashed line). Crust-mantle boundary (Moho) is determined

30

557 *from the Bouguer anomaly field applying gravity inverse calculations using Lithoflex*
558 *software package (www.lithoflex.org, Braitenberg et al. 2007; Wienecke et al. 2007).*
559 *Individual centers at the active arc front are indicated with white triangles (from*
560 *Siebert and Simkin, 2002). OSL is the Ojos del Salado-San Buenaventura volcanic*
561 *lineament. Chilean trench and coastline are indicated by black and white triangles*
562 *respectively. **b.** Series of east-west cross sections representing the Wadati-Benioff zone*
563 *(indicated in blue) from Mulcahy et al. (2014) showing how the flat slab is well*
564 *developed south of 28.25°S, while at 27°S indicates a shallow subduction setting.*

565

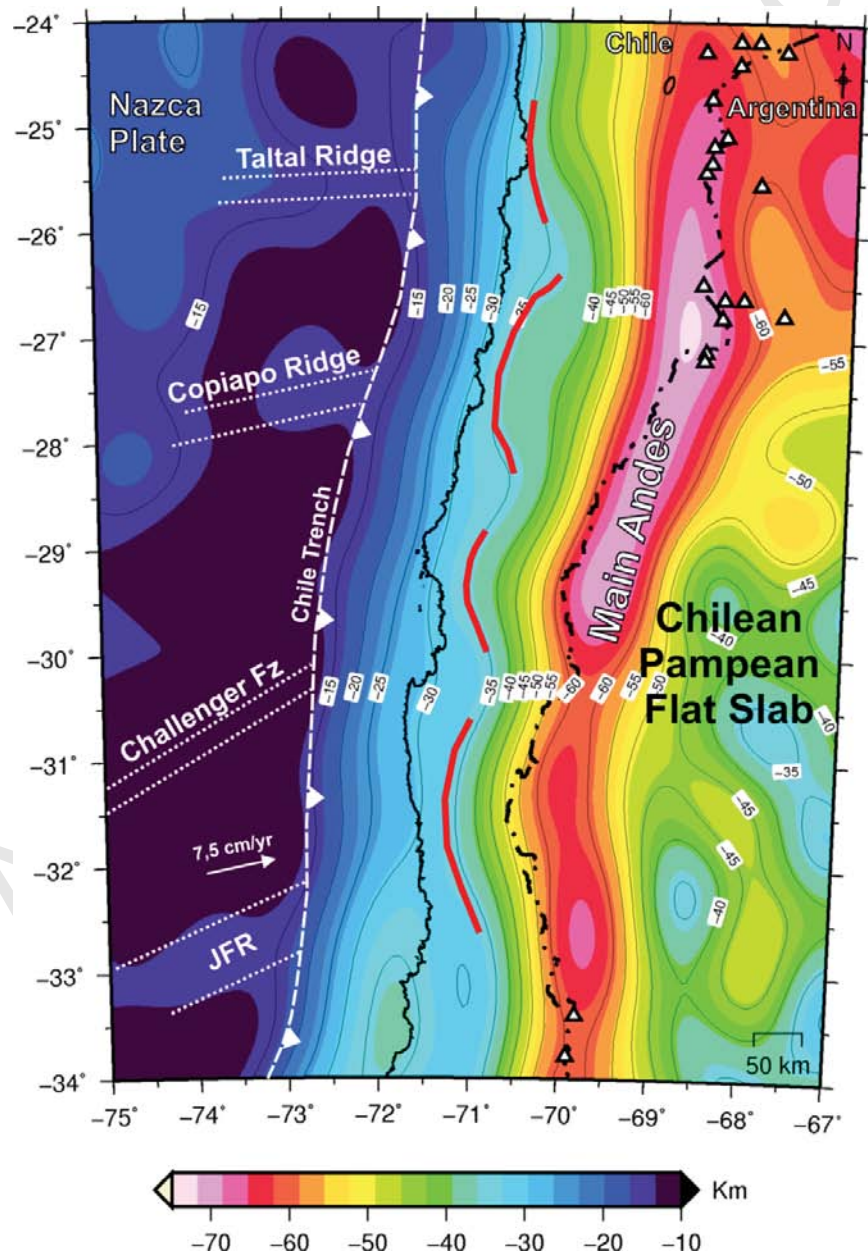
566 **5. DISCUSSION AND CONCLUDING REMARKS**

567 Flat subduction settings modify the pattern of seismicity, volcanism and deformation in
568 the overriding plate. Results from analogue modeling (Martinod et al. 2013) show that
569 indentation of an overriding plate during ridge subduction is usually accommodated by
570 a complex pattern of structures that include trench-perpendicular strike-slip faults. In
571 particular, the Ojos del Salado-San Buenaventura volcanic lineament represents an
572 anomaly in the pattern of arc-retroarc volcanism at the northern extreme of the Chilean-
573 Pampean flat slab (Figs. 2, 3, A1 and A2). These volcanic centers are controlled by
574 ENE-oriented regional neotectonic structures with dip and strike-slip strain components
575 that project into the retroarc zone transversally to the arc front. ~~Results from analogue~~
576 ~~modeling (Martinod et al. 2013) show that indentation of the overriding plate during~~
577 ~~ridge subduction is usually accommodated by trench-perpendicular strike-slip faults.~~
578 Locally, this volcanic lineament coincides with a strong deflection in the fabric of the
579 Andean deformation visualized from the gravity analyses (Fig. 5b). The observed

580 rotation in the general strike of the NNE “Andean” pattern (Fig. 4a, b) of the Ga and Tzz
581 anomalies (obtained from EGM2008 model) is consistent with a strong deflection in the
582 Andean drainage divide to ENE in this region; imposed by the aligned summit of
583 volcanic centers and neotectonic structures associated with the Ojos del Salado-San
584 Buenaventura volcanic chain. More specifically the volcanic centers (yellow circles of
585 Fig. 4a) coincide with positive values of the Tzz signal (relatively denser bodies are
586 related to a positive Tzz) and with relatively higher values in the Ga , indicating higher
587 density materials (or a positive density contrast).

588 These anomalous neotectonic pattern and volcanic alignment across the Andes at about
589 27.5°S could be spatially linked to the collision of the Copiapó aseismic ridge. The
590 strike ~~Direction~~ of the Copiapó ridge parallels the plate convergence direction between
591 the Nazca and South American plates (78.1° azimuth NE in our study area, Kendrick et
592 al. 2003), determining a stationary collisional point (Fig. 5b), similarly to the point of
593 inception of the Juan Fernández ridge to the south in the last 10 My (Yáñez et al. 2001).
594 Therefore, the deformational imprint over the upper plate would be expected to have
595 affected a discrete area ~~zone~~, producing an ENE-localized deformational zone,
596 explaining the local deflection of gravity anomalies (Fig. 11). Similarly, neotectonic
597 deformation acquires, east of the arc front, a predominant E-NE orientation along the
598 San Buenaventura volcanic alignment (Seggiaro et al. 2000; Zhou et al. 2014). This
599 deformation is clearly depicted by the Tzz (from EGM2008 model) that constitutes a
600 gravity derivative that highlights superficial density anomalies. Thus, it allows the
601 detection of the edge of geological structures and to distinguish the signal due to a
602 smaller shallow density variation from an extensive deeper mass. The spectral power of
603 Tzz signal is pushed to higher frequencies, resulting in a signal more focalized to the

604 source than the *Ga* (Li 2001), being the last more sensitive to regional signals and
 605 deeper sources (Alvarez et al. 2012). The long wavelength signal from GOCE model
 606 also shows an eastward deflected *Ga*, suggesting that the upper crust deformation
 607 showed by the *Tzz* could be related to deeper sources; connected by oblique-to-the-
 608 Andes intracrustal structures acting as magmatic paths for arc and retroarc volcanism,
 609 promoting long volcanic alignments as the OSL-San Buenaventura.



610

33

611 **Figure 10.** *Moho depth determined from inversion of GOCE data between 25°S and*
612 *34°S. Note from the forearc to the arc regions an anomalous Moho thickening at the*
613 *sites of inception of four high oceanic features (anomalous buoyant oceanic crust)*
614 *~~aseismic ridges~~ in the Chilean-Peruvian trench, and in particular at the sites of*
615 *inception of the Juan Fernández and Copiapó ridges.*

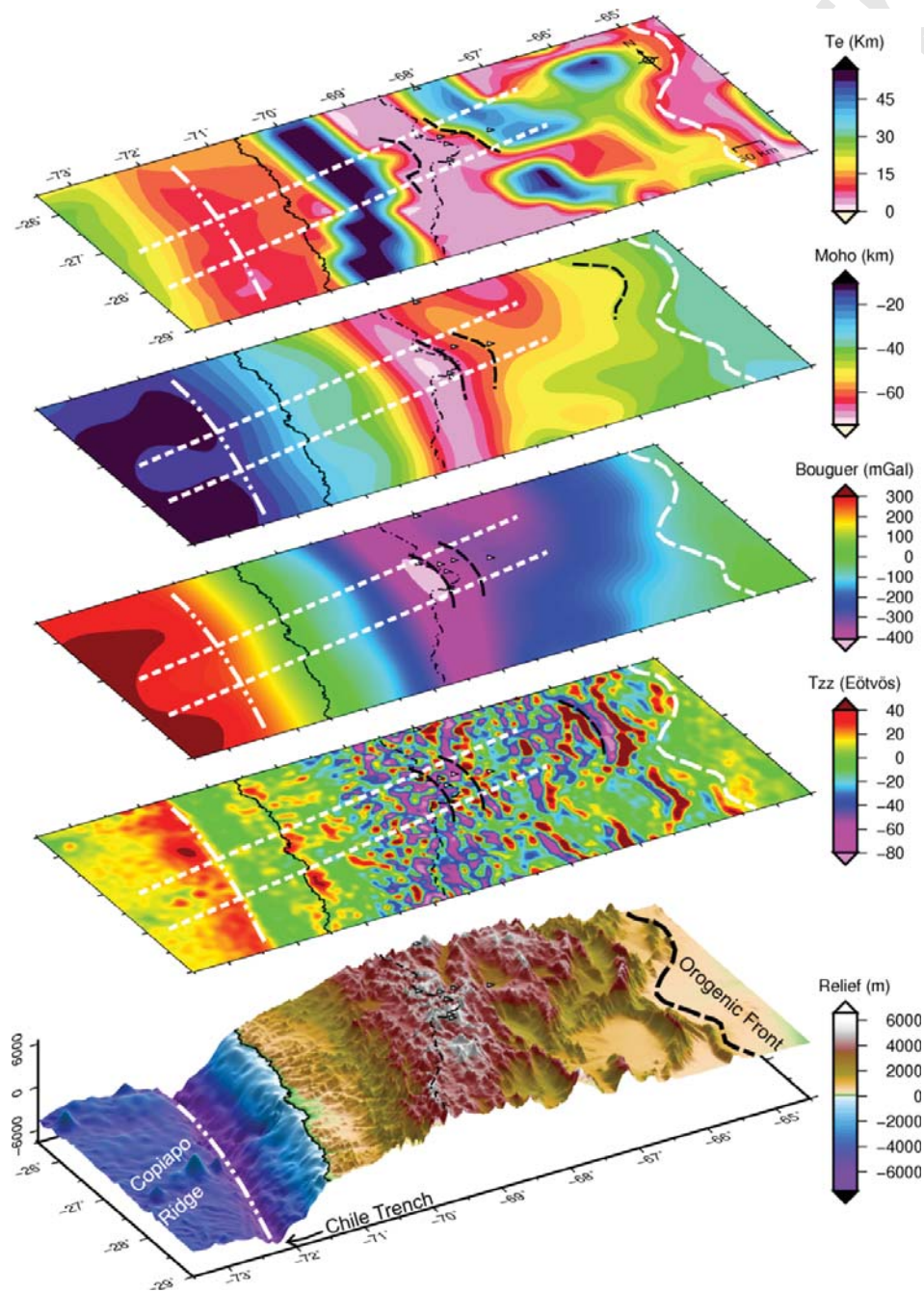
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617 ~~A maximum crustal thickness at the arc zone is inferred at the zone of inception of the~~
618 ~~Copiapó ridge from inversion of gravity data (Fig. 6b). Similarly, at the forearc region a~~
619 ~~relative crustal thickening respect to neighbor segments is inferred from inversion of~~
620 ~~gravity data (Fig. 6b). From inversion of gravity data, we obtained a maximum crustal~~
621 ~~thickness at the arc and forearc zones, coincident with the path of subduction of the~~
622 ~~Copiapó ridge (Fig. 6b). In fact, crustal thickening at the forearc zone can be checked in~~
623 ~~this model throughout each zone of inception of a high oceanic feature that impact~~
624 ~~against the Chilean subduction margin such as the JFR, Challenger Fz, Copiapó and~~
625 ~~Taltal ridges (Fig. 10). Lower crustal thickening- Crustal thickening at the forearc and~~
626 ~~arc zones, as well as extensional and strike slip deformation affecting the upper crust~~
627 ~~and related controlling volcanism at the Ojos del Salado-San Buenaventura volcanic~~
628 ~~lineament could be interpreted as a consequence of higher coupling, localized~~
629 ~~shortening, underplated volcanic materials and regional upward doming. Upper crustal~~
630 ~~uplift is evidenced by high positive T_{zz} -EGM2008 values over the volcanic centers,~~
631 ~~while the thickening of the lower crust is detected by negative G_a -GOCE at the site of~~
632 ~~ridge subduction over the flat slab zone.~~

633 On a more general perspective, flat to shallow subduction zones around the world can
634 be separated in two general groups based on their size. Small zones such as the one
635 associated with the collision of the Carnegie ridge in Ecuador and the one related to the
636 subduction of the Juan de Fuca plate next to Vancouver island in Canada, and large flat
637 slabs such as the Peruvian and the Chilean-Pampean along the central South American
638 subduction margin (Gutscher et al. 2000). In particular the Peruvian and the Chilean-
639 Pampean flat slabs at both sides of the Altiplano-Puna plateau are the most important in
640 terms of size and associated crustal deformation. Particularly, the Peruvian flat slab has
641 been related to the collision of two aseismic ridges, one completely subducted beneath
642 the South American plate (Gutscher 1999a; Rosenbaum et al. 2003; Rouse et al. 2003).
643 Contrastingly, the Chilean-Pampean flat slab has been assigned to the collision of a
644 single aseismic ridge, explaining the extent of this zone by a broken geometry of the
645 subducted ridge beneath the South American plate (Anderson et al. 2007; Yañez et al.
646 2001).

647 The Benioff zone geometry at the northern region of the Chilean-Pampean flat slab (Fig.
648 7) presents a strong change in the slab dip from normal to shallow. Recent works based
649 on local seismic networks (Mulcahy et al. 2014) show that the northern and southern
650 terminations of the flat slab are more abrupt than previously determined, resulting in a
651 flat slab configuration rather symmetrical in map view. Additionally, both edges of the
652 flat slab coincide with the sites of inception of the Copiapó and Juan Fernandez ridges
653 respectively. In this sense, the development of the Chilean-Pampean flat slab could be a
654 function of two simultaneous ridge collisions, similarly to the Peruvian flat slab to the
655 north (Figs. 11 and 12). Finally, this study exemplifies how the Earth Gravity Model
656 EGM2008, used in combination with the newest satellite GOCE data, can be used to

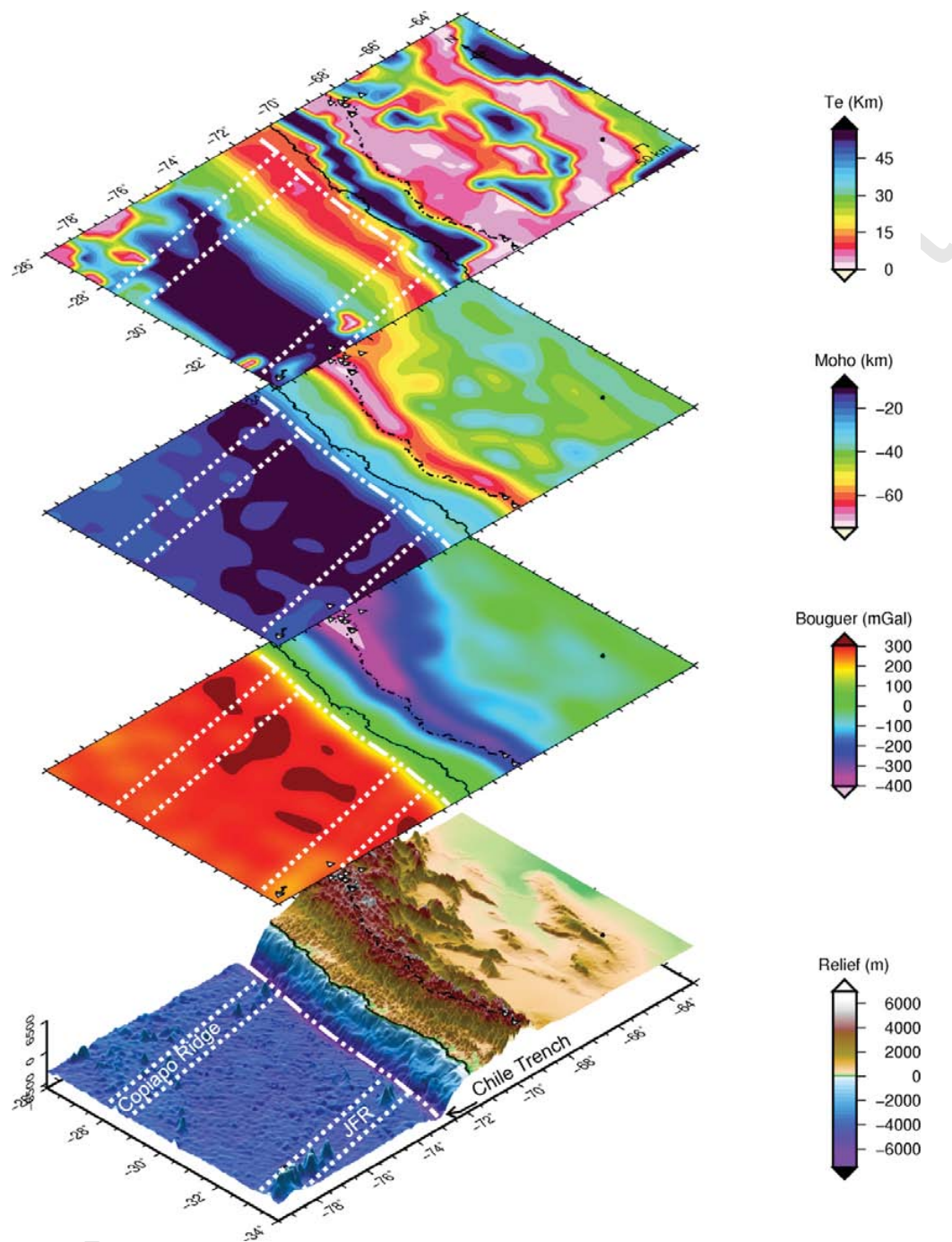
657 track the fate of certain subducted bathymetric anomalies beneath the upper plate of a
 658 subduction zone, connecting anomalously thickened ~~subducted~~ oceanic crust,
 659 deformational effects, T_e anomalies, seismic patterns and variable crustal thickness in
 660 the overriding plate, ~~and variable crustal thickness.~~



661

36

662 **Figure 11.** *Relation between the elastic thickness, Moho, Bouguer anomaly, T_{zz} and*
663 *topography at the area of inception of the Copiapó ridge beneath the South American*
664 *plate. Note the relation between the extrapolated ridge path and the change in*
665 *orientation and truncation of the anomalies highlighted in the gradient signal (T_{zz}), the*
666 *eastward deflection observed in Bouguer anomaly, a deeper Moho topography and*
667 *higher T_e values.*



668

669 **Figure 12.** *Two ridge collisions and crustal structure at the Chilean-Pampean flat slab.*

670 *Note similar rigidity structure and Moho amplitudes at both sites of ridge inception.*

671 *While the southern ridge collision (Juan Fernández ridge) is linked to a flat slab*

672 *configuration without arc magmatism, the northern collision coincides with a shallow*

38

673 *subduction configuration that passes transitionally to a normal subduction setting in the*
674 *north, with a volcanic arc migrating eastwards at the site of inception of the Copiapó*
675 *ridge beneath the South American continent.*

676

677 **ACKNOWLEDGMENTS:** Authors acknowledge the use of the GMT-mapping
678 software of Wessel and Smith (1998). We greatly acknowledge S.M. Kay for
679 discussion. The authors would like to thank the Ministerio de Ciencia y Técnica –
680 Agencia de Promoción Científica y Tecnológica, PICT07–1903, Agenzia Spaziale
681 Italiana for the GOCE-Italy Project, the Ministero dell’Istruzione, dell’Universita’ e
682 della Ricerca (MIUR) under project PRIN, 2008CR4455_003 for financial support, and
683 ESA for granting of AO_GOCE_proposal_4323_Braitenberg.

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1249 faulting in the Pasto Ventura region of the Southern Puna Plateau, NW Argentina.
1250 Tectonics. In Press.

1251 **APPENDIX A: Magmatic evolution of the southern Puna plateau**

- 1252 Over flat subduction settings, the volcanic arc migrates and expands through the
1253 foreland area and consequently thermal crustal structure changes, producing the

1254 shallowing of brittle-ductile transitions that favor shortening of the upper crust
1255 (see James and Sacks, 1999 and Kay and Coira 2009). In particular, the Chilean-
1256 Pampean flat subduction zone, initiated ~17 My ago, has been related to the
1257 exhumation of the foreland area through the activity of decollements inaugurated
1258 since the arc foreland excursion (see Ramos et al. (2002), for an extended
1259 discussion).

1260 The volcanic centers within the Ojos del Salado-San Buenaventura volcanic
1261 lineament are affected and controlled by normal and right lateral neotectonic
1262 structures, defining a general ENE structural alignment (Fig. 3) (Kay et al. 2008;
1263 Marrett et al. 1994; Mpodozis et al. 1996; Seggiaro et al. 2000; Zhou et al.
1264 2014). Most of the volcanic centers controlled by these structures are Pleistocene
1265 in age, although some show a more prolonged activity during the Holocene
1266 (Baker et al. 1987; Kraemer et al. 1999; Montero López, 2009; Mpodozis et al.
1267 1996; Risse et al. 2008; Seggiaro et al. 2000; Viramonte et al. 2008). Based on
1268 radiometric ages compiled along the Ojos del Salado-San Buenaventura volcanic
1269 lineament, four magmatic stages (>8My, 8-4My, 4-2My and <1My) can be
1270 recognized with distinctive distributions across the arc and retroarc zones (Figs.
1271 A1 and A2).

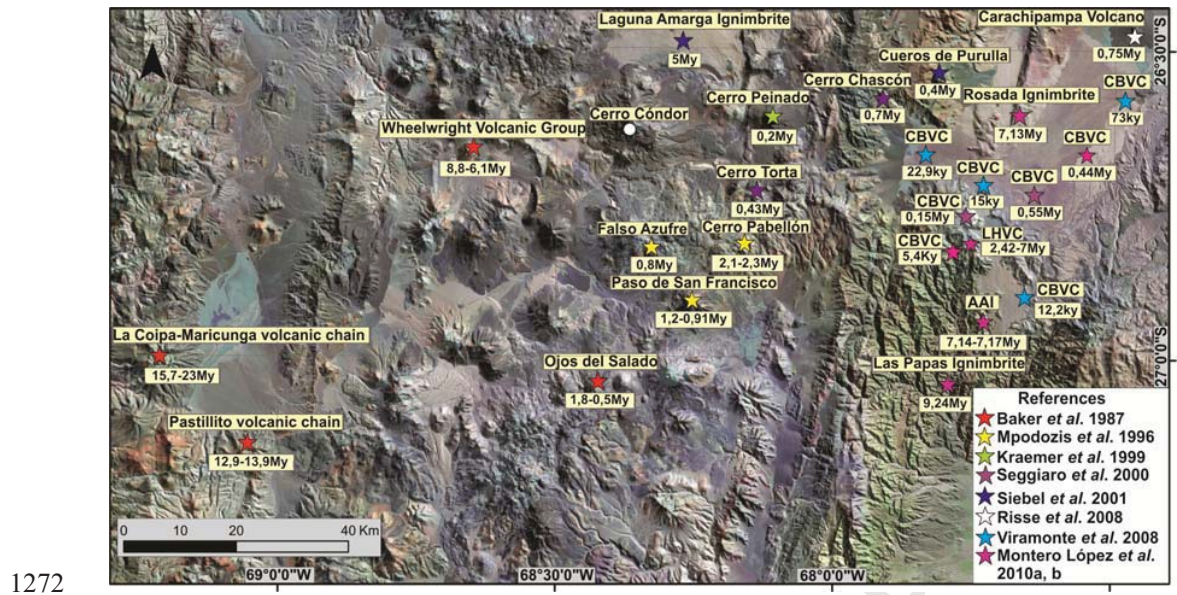
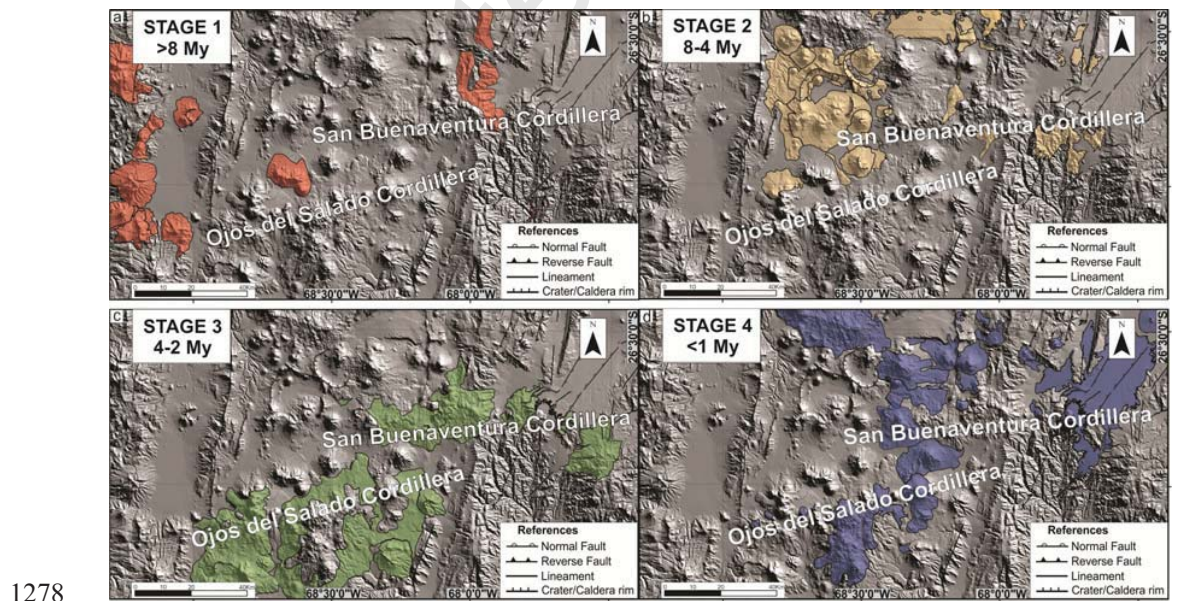


Figure A1. Radiometric ages in volcanic centers along the Ojos del Salado-San Buenaventura volcanic lineament from the arc front to the eastern retroarc area, draped on top of a TM image (radiometric ages were compiled from Baker et al. 1987; Kraemer et al. 1999; Montero López, 2009; Mpodozis et al. 1996, Seggiaro et al. 2000; Risse et al. 2008; Viramonte et al. 2008).



1279 **Figure A2.** *Spatial and temporal distribution of volcanism in the Altiplano-*
1280 *Puna Plateau from the arc front to the retroarc area through the Ojos del*
1281 *Salado-San Buenaventura volcanic lineament (Coira y Kay 1993; Kay and*
1282 *Coira 2009; Kay et al. 2010; Salisbury et al. 2011; Trumbull et al. 2006).*

1283

1284 An older stage (>8My) is mainly represented on the Chilean Andean slope by La
1285 Coipa-Maricunga and Pastillito volcanic associations (Baker et al. 1987),
1286 represented by stratovolcanoes and domes associated with ignimbrite deposits
1287 with a general northward trend. In particular, the Pastillito volcanic association
1288 (12.9-13.9My) is developed 15-20 km east of previous volcanism in the region.
1289 At the end of this stage, the eastward migration of the volcanism is more evident
1290 (Kay and Coira, 2009), with middle to late Miocene volcanic centers located in
1291 the San Buenaventura volcanic lineament (Montero López et al. 2010a, 2011;
1292 Seggiaro et al. 2000). The 8-4 My stage comprises mainly riodacitic calderas
1293 with associated ignimbrites (e.g. Wheelwright caldera, Laguna Amarga
1294 Ignimbrite and Rosada Ignimbrite) and backarc andesitic volcanism of La
1295 Hoyada Volcanic Complex (Montero López et al. 2010b). The 4-2 My stage
1296 marks the onset of a different tectonic control with volcanic centers aligned
1297 through a NE trend (Baker et al. 1987). In this stage, the Peñas Blancas Volcanic
1298 group and the main volume of the San Buenaventura volcanic lineament was
1299 built. The last stage, <1My, is represented by the emplacement of a series of
1300 stratovolcanoes such as San Francisco, IncaHuasi, Falso Azufre, Ojos del Salado
1301 and Cónдор, and the Cerro Blanco Volcanic Complex towards the back arc
1302 region (Arnosio et al. 2008; Montero López et al. 2010a; Viramonte et al. 2008).

1303

1304 **APPENDIX B: Statistical comparison between GOCE and EGM2008 models**

1305 Even though recent satellite gravity missions have gained an extraordinary
1306 improvement in the global mapping of the Earth gravity field, the models derived solely
1307 from data of these missions (e.g. Pail et al. 2011) present a lower spatial resolution than
1308 mixed satellite-terrestrial global models like EGM2008 (Pavlis et al. 2008). GOCE data
1309 were obtained through acceleration and gravity gradient measurements, and also
1310 through orbit monitoring. For degrees greater than $N=120$, EGM2008 relies entirely on
1311 terrestrial data, so GOCE only models, as the GO_CONS_GCF_2_TIM_R4 with
1312 $N=250$ (Pail et al. 2011), are a remarkably important independent quality assessment
1313 tool to control the quality of the terrestrial data entering in EGM2008 (Braitenberg et al.
1314 2011a; Alvarez et al. 2012, 2014). A comparison analysis up to the maximum degree
1315 ($N=250$) of the GOCE model is a simple way to evaluate the contribution of the
1316 terrestrial data between $N=120$ and $N=250$ entering EGM2008. If the residual is small
1317 ground-based data represent accurately the field up to degree $N=250$ relying on the
1318 correctness for higher orders (see Braitenberg et al. 2011b; Alvarez et al. 2012). The
1319 errors of the original terrestrial data affect the errors of the EGM2008 values up to
1320 $N=250$, because the spherical harmonic expansion can be seen as an averaging process.
1321 The standard deviations between GOCE and EGM2008 thus represent varying quality
1322 of the original terrestrial data, because the quality of the GOCE data is locally
1323 homogeneous. Where the standard deviations are small, the original data must have
1324 been accurate or otherwise the same downscaled values and a small standard deviation
1325 would have been only obtained by chance (See Braitenberg et al. (2011a) and Alvarez et
1326 al. (2012) for a more detailed explanation). Yi and Rummel (2014) made a comparison

1327 of GOCE models with EGM2008 and found that the agreement between EGM2008 and
 1328 the GOCE-models up to degree and order 200 is good, with a global (excluding the
 1329 polar gaps of GOCE orbits, throughout) geoid RMS-difference of 11 cm in the ocean
 1330 areas and between 8-20 cm in the continental areas. Therefore GOCE is a remarkably
 1331 important independent quality assessment tool for EGM2008, especially in those areas
 1332 where no precise terrestrial data are available in the EGM2008 model as is the case of
 1333 vast parts of South America. Here, substantial differences are expected, especially in the
 1334 Andes region where the topography is high and rough in many areas. An assessment for
 1335 the comparison of terrestrial observations and EGM2008 and GOCE was made by
 1336 Bomfim et al. (2013).

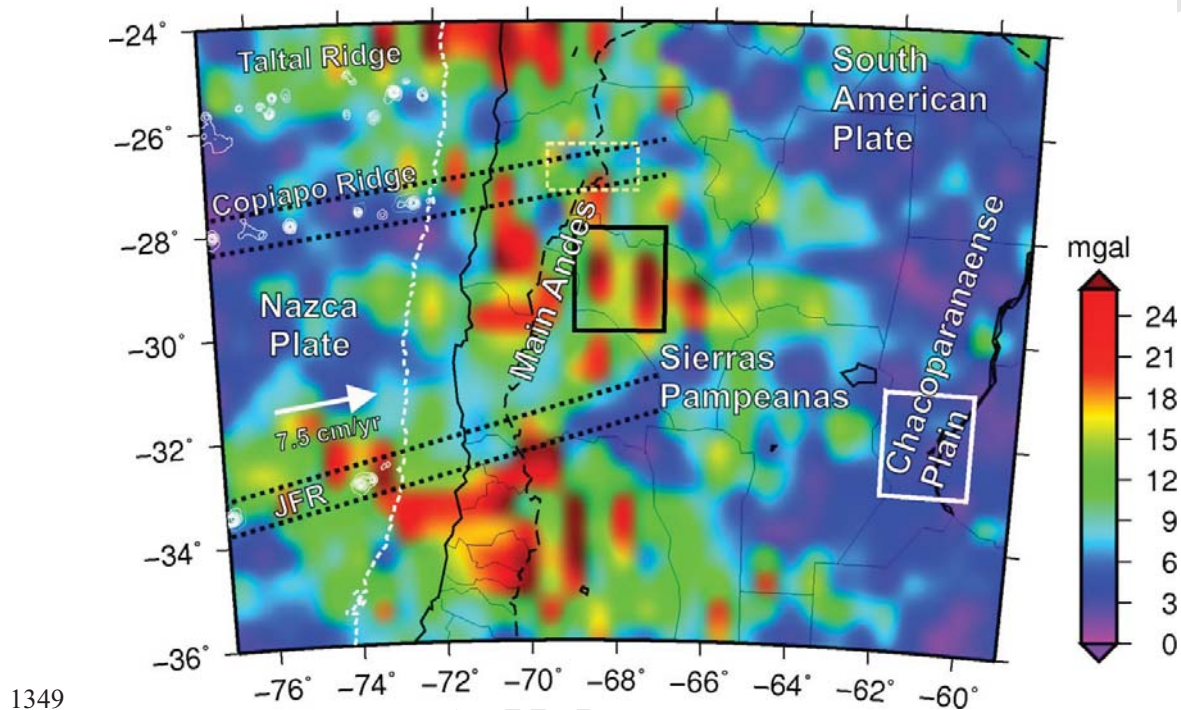
1337 The absolute residual (Fig. B1) between the gravity anomaly derived from GOCE
 1338 model (Pail et al. 2011) and the gravity anomaly derived from the EGM2008 model
 1339 (Pavlis et al. 2008, 2012) shows that the fields are in partial agreement. Statistical
 1340 parameters for the difference between the two fields are shown on Table B1. A high-
 1341 quality region is compared with a low-quality region in terms of the residual histogram.
 1342 An area with degraded quality, corresponding to the 2° x 2° black square in Fig. B1, is
 1343 compared to a square of equal size (white) of relatively high quality.

<i>Average difference</i>	<i>0.0772 mGal</i>
<i>Standard Deviation</i>	<i>12.341 mGal</i>
<i>Maximal value of difference</i>	<i>62.021 mGal</i>

1344 **Table B1:** *Statistical parameters for the difference.*

1345

1346 The histograms of the residuals (Fig. B2) illustrate the higher values for the black
 1347 square and a limited error (± 10 mGal) for the white square. Instead, the black square,
 1348 presents a high error ($-45/+35$ mGal) with a uniform distribution.



1349

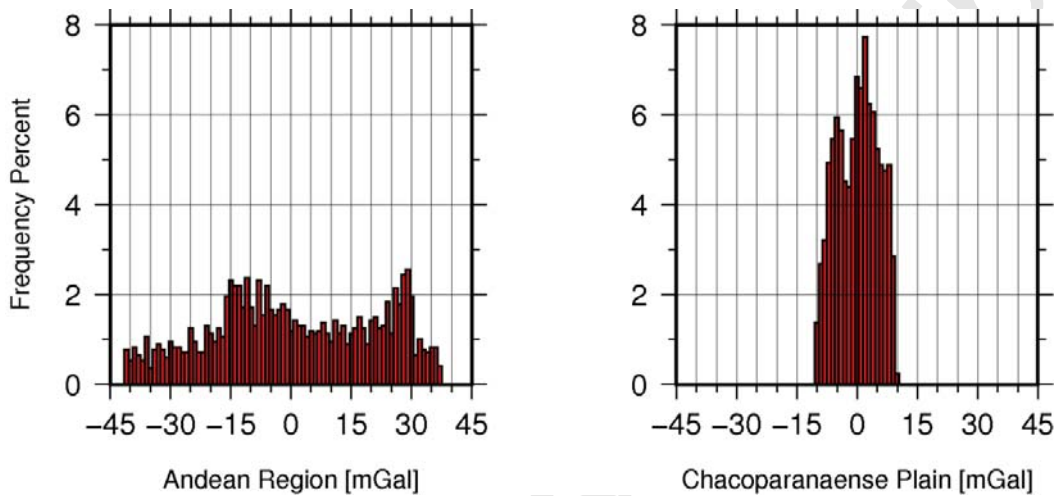
1350 **Figure B1.** Absolute difference between the two fields. The black square shows the area
 1351 with erroneous data, while the white square shows the area with good data agreement.
 1352 National borders: dashed black line; coastal borders: solid black line. The yellow
 1353 dashed rectangle represents the approximate area where the deflection of the Andean
 1354 structures was detected by means of the G_a and T_{zz} from EGM2008 model (Figs. 4a, b
 1355 and 5b).

1356

1357 The absolute difference between both fields shows a range of variation between
 1358 approximately $+6$ mGal and $+24$ mGal in the area where the Andean structures were

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1359 deflected (Figs. 4a, b and 5b) and detected by means of the Ga and Tzz from EGM2008
 1360 model (yellow dashed rectangle in fig B1). The gravity signal that shows such pattern
 1361 reaching less than -450 mGal (Fig. 4b), thus giving a signal to noise ratio of
 1362 approximately $SNR \approx 24 \text{ mGal} / 450 \text{ mGal} = 0.05333$.

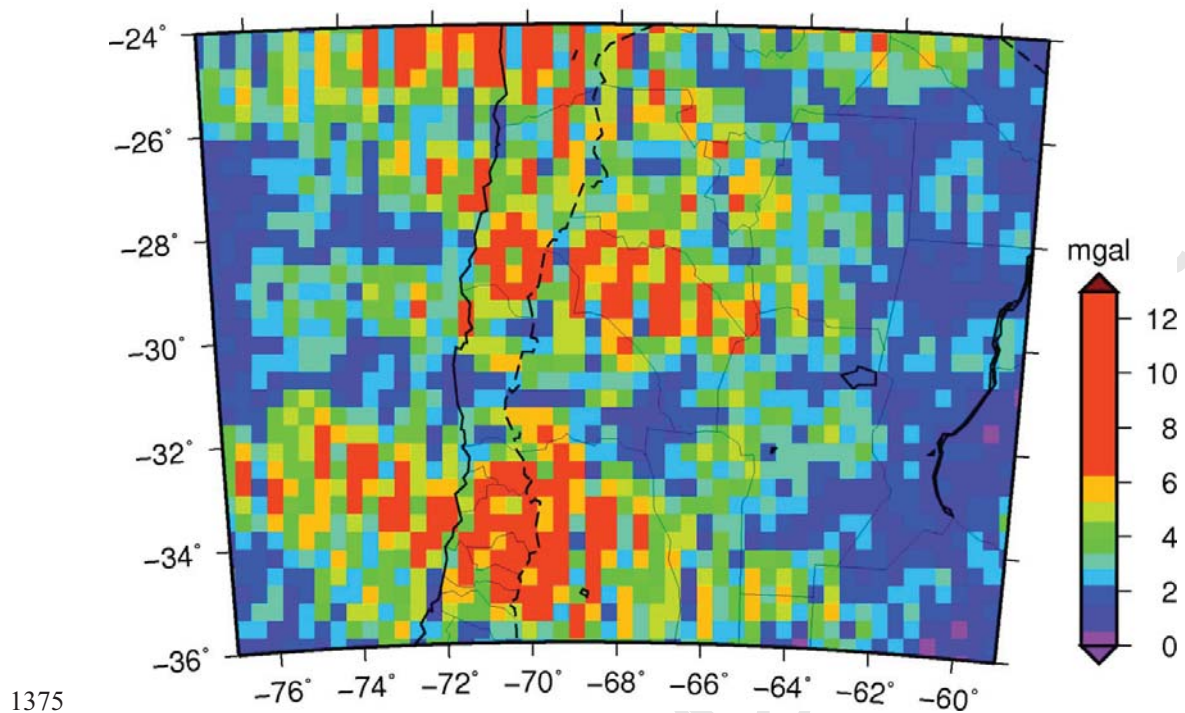


1363

1364 **Figure B2.** Histogram of the residual gravity anomaly between EGM2008 and GOCE
 1365 (up to degree and order $N=250$). Left (Good tile): white square of Fig. B1. Right (Bad
 1366 tile): black square of Fig. B1.

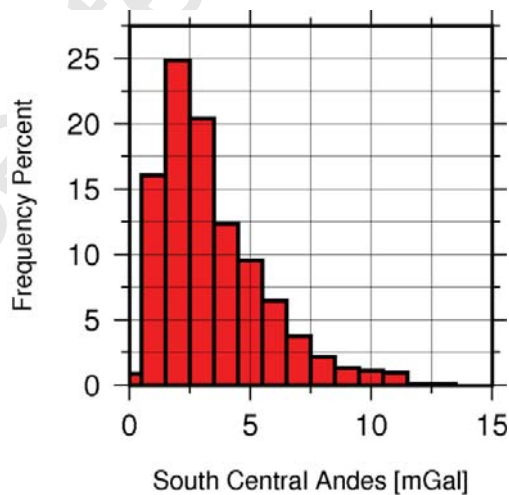
1367

1368 As a statistical measure of EGM2008 quality, the root mean square (rms) deviation
 1369 from the mean was calculated, on sliding windows of $20' \times 20'$. The result is shown on
 1370 Fig. B3. The most frequent value of the rms deviation is 2 mGal (Fig. B4). The regions
 1371 where the EGM2008 present higher differences of the rms difference, when compared
 1372 to GOCE, reach up to 12.5 mGal. The locations where the terrestrial data present the
 1373 main differences reflected up to 12 mGal. Approximately the 85.5% of the rms
 1374 deviation is below 5 mGal.



1376 **Figure B3.** Root mean square of the gravity residual on 20' x 20' tiles.

1377



1378

1379 **Figure B4.** Histogram of the rms deviations on 0.5° x 0.5° tiles.

1380

1381

1382 **APPENDIX C: Calculation of the flexural strength of the lithosphere**

1383 To accomplish the inverse modeling of the flexural rigidity, we used Lithoflex software
1384 package (www.lithoflex.org). Isostatic modeling adopts the isostatic thin plate flexure
1385 model (e.g. Watts, 2001) and the use of a newly derived analytical solution for the 4th
1386 order differential equation that describes the flexure of a thin plate (concept introduced
1387 by Vening-Meinesz in 1939). This allows the analytical calculation of the deflection of
1388 a thin plate for any irregular shape of the topography (see Wienecke 2006 and
1389 references therein for a more detailed discussion). To estimate the elastic properties of
1390 the plate for a known load, a crustal load and the crust-mantle interface to be used as a
1391 reference surface are needed (Wienecke, 2006). The load acting on the crust is
1392 constituted by the combination of the overlying topography plus a density model
1393 (Braitenberg et al. 2007). A density variation within the crust represents a variation in
1394 the load, and must be reflected in the isostatic response (Ebbing et al. 2007). Long
1395 wavelength information of the gravity field mainly corresponds to the crust/mantle
1396 density contrast, but sedimentary basins can also produce a long wavelength signal thus
1397 influencing the correct gravity crust mantle interface estimation by gravity inversion
1398 process (Wienecke 2006). Therefore, the gravity effect of sediments was calculated in
1399 order to reduce the gravity data.

1400 The topographic load or equivalent topography was calculated using
1401 topographic/bathymetric data from ETOPO1 (Amante and Eakins, 2009). The densities
1402 used for calculation were $1,030 \text{ kg/m}^3$ for water and $2,800 \text{ kg/m}^3$ for the crust. This
1403 method requires two input parameters: the density contrast and reference depth.
1404 Standard parameters such as normal crust thickness $T_n = 35 \text{ km}$, and a crust mantle

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1405 density contrast of -400 kg/m^3 were used. The Bouguer anomaly was calculated using
 1406 the GOCE satellite data (GO_CONS_GCF_2_TIM_R4, [http://icgem.gfz-](http://icgem.gfz-potsdam.de/ICGEM/)
 1407 [potsdam.de/ICGEM/](http://icgem.gfz-potsdam.de/ICGEM/), Pail et al. 2011) up to degree/order $N=250$ and reduced by the
 1408 effect of sediments.

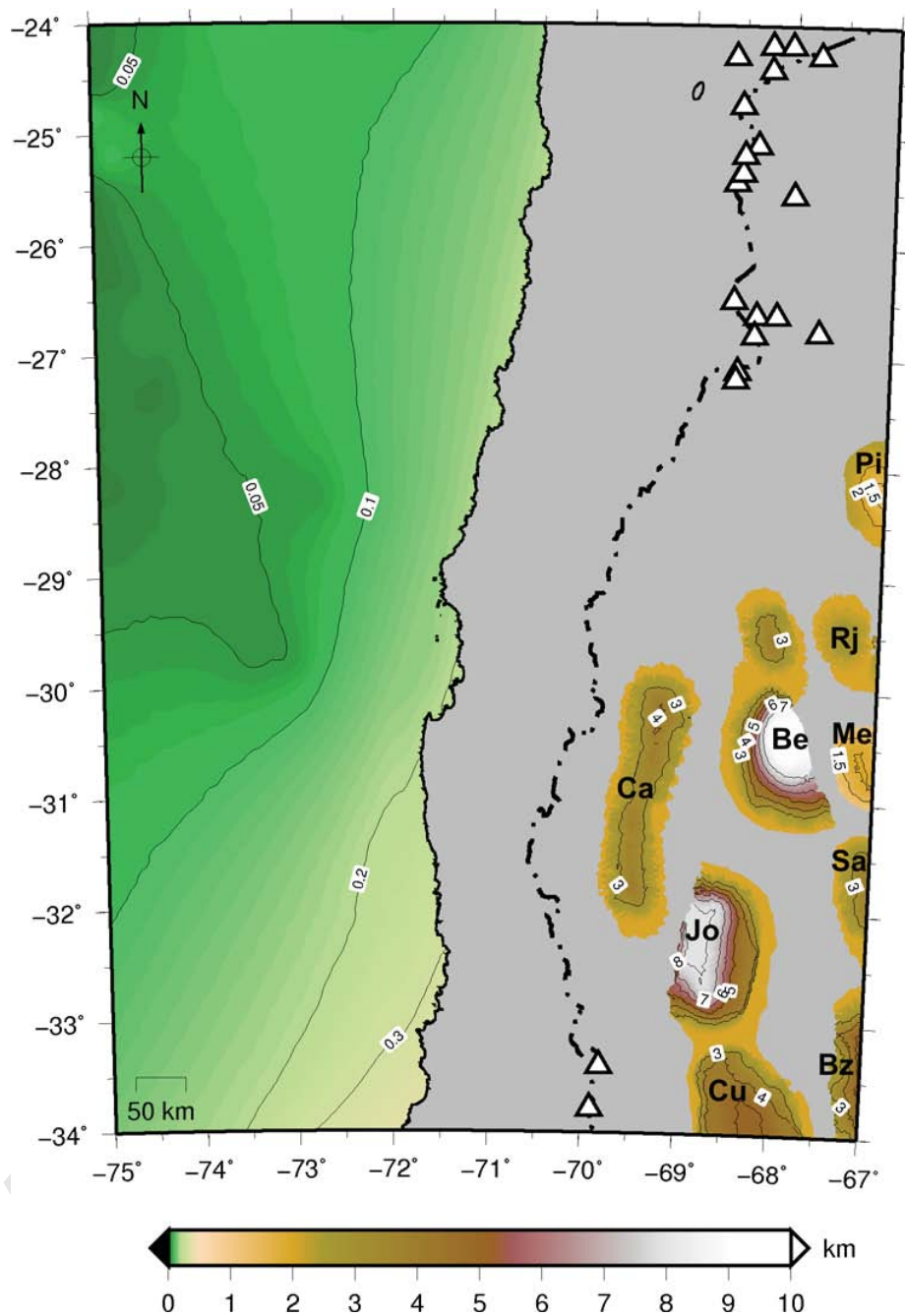
1409 Forward calculation of the gravity effect of the basin fillings (Fig. C1) was done taking
 1410 into account a linear variation of density with depth (see Braitenberg et al. 2007). To
 1411 perform this calculation we defined a two-layer reference model of the continental crust
 1412 with an upper crustal density of $2,700 \text{ kg/m}^3$ and a lower crustal density of $2,900 \text{ kg/m}^3$.
 1413 To perform this operation we used the bathymetry from ETOPO1 (Amante and Eakins,
 1414 2009) and off-shore sediment thicknesses from Whittaker et al. (2013). On-shore basins
 1415 were modeled using depths to the top of the basement calculated from gravimetric
 1416 studies and seismic lines of Yacimientos Petroliferos Fiscales (YPF), Texaco, Repsol
 1417 YPF, YPF S.A. and OIL M&S, and from Barredo et al. (2008); Fernandez Seveso and
 1418 Tankard (1995); Kokogian et al. (1993), Milana and Alcober (1994), Miranda and
 1419 Robles (2002), Rosello et al. (2005) and Gimenez et al. (2009). The correction amounts
 1420 up to -45 mGal for the main on-shore basins and up to a few mGal for oceanic
 1421 sediments reaching their maximum over the Chilean trench.

Masses above sea level	ρ_s	$2,670 \text{ kg/m}^3$
Upper crustal density	ρ_{uc}	$2,700 \text{ kg/m}^3$
Lower crustal density	ρ_{lc}	$2,900 \text{ kg/m}^3$
Upper mantle density	ρ_m	$3,300 \text{ kg/m}^3$
Young modulus	E	10^{11} N/m^2
Poisson ratio	Σ	0.25

1422

Table C1: *Parameters used in the flexural modeling.*

1423

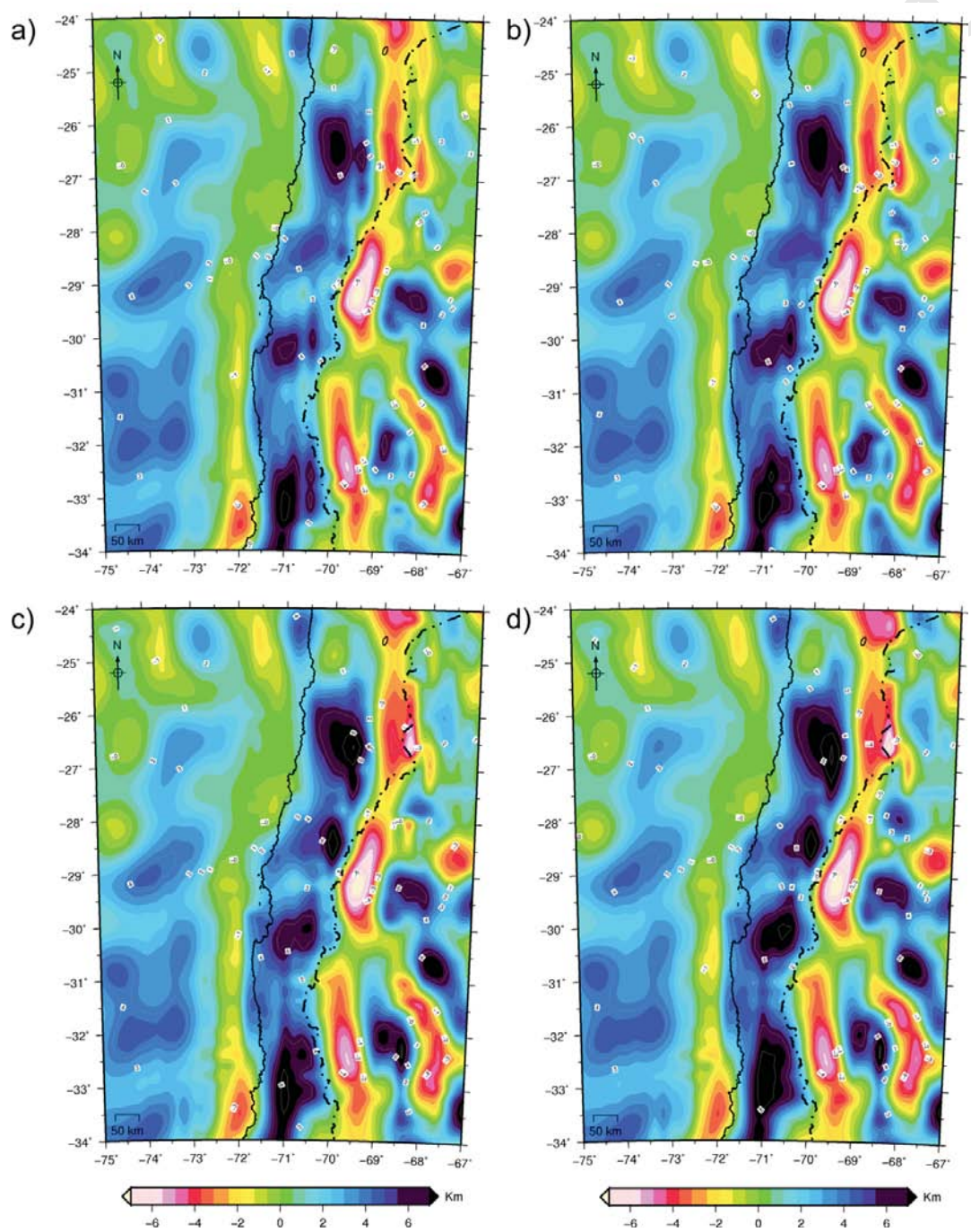


1424

1425 **Figure C1.** *Sediment thicknesses used to reduce the gravity data. Offshore sediment*
 1426 *thickness is from Whittaker et al. (2013). The basins over the South American plate*
 1427 *were approximated using gravity databases and basement depths from seismic lines*

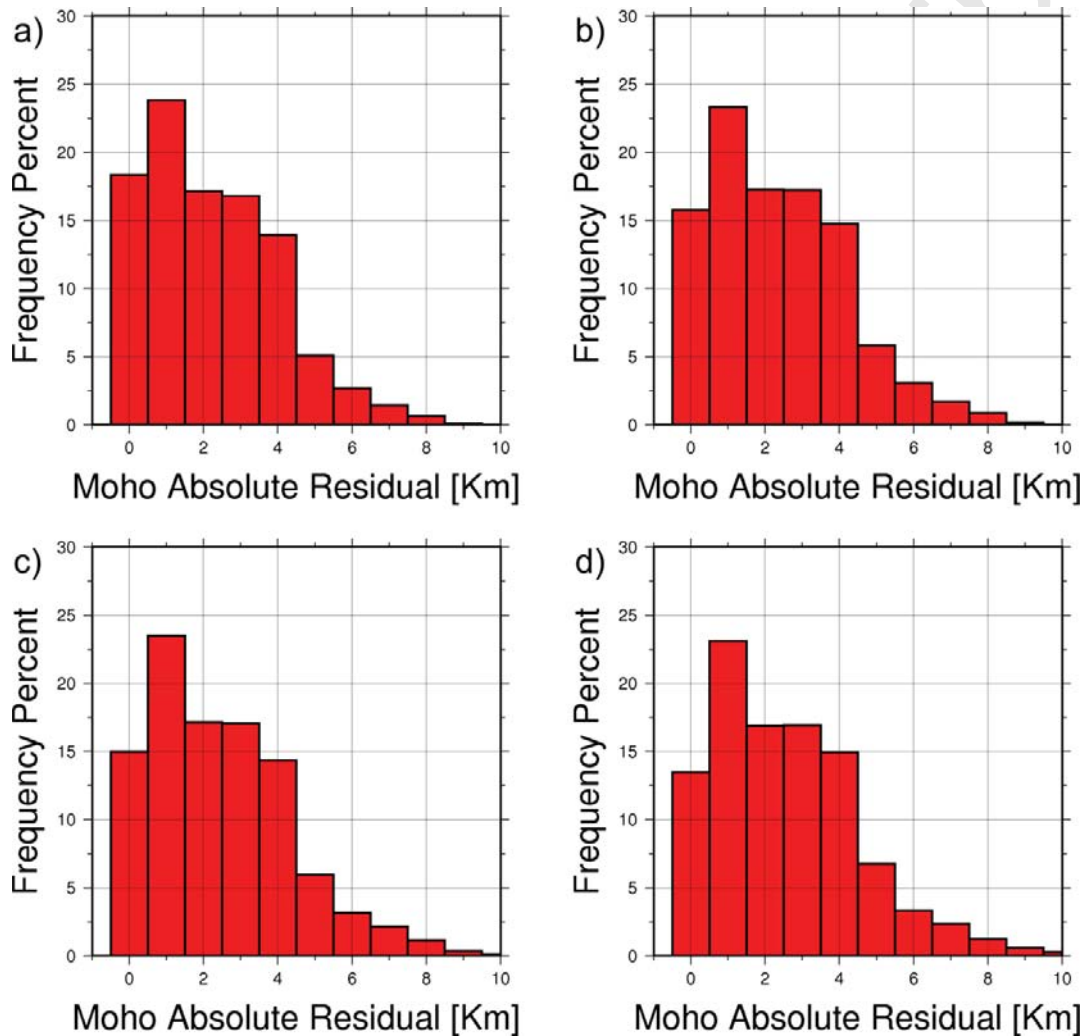
77

1428 (YPF S.A. unpublished report). Sedimentary basins: Be, Bermejo; Jo, Jocoli; Sa,
1429 Salinas; Cu, Cuyana; Bz, Beazley, Ca, Calingasta; Pi, Pipanaco; Rj, La Rioja; Me,
1430 Medanos.



1431

1432 **Figure C2.** Residual obtained by subtracting the crust–mantle interface obtained by
 1433 gravity inversion of the sediment corrected Bouguer anomaly minus the crust–mantle
 1434 interface obtained from the flexural model. **a)** Window size 60x60km. **b)** Window size
 1435 70x70km. **c)** Window size 80x80km. **a)** Window size 90x90km.



1436

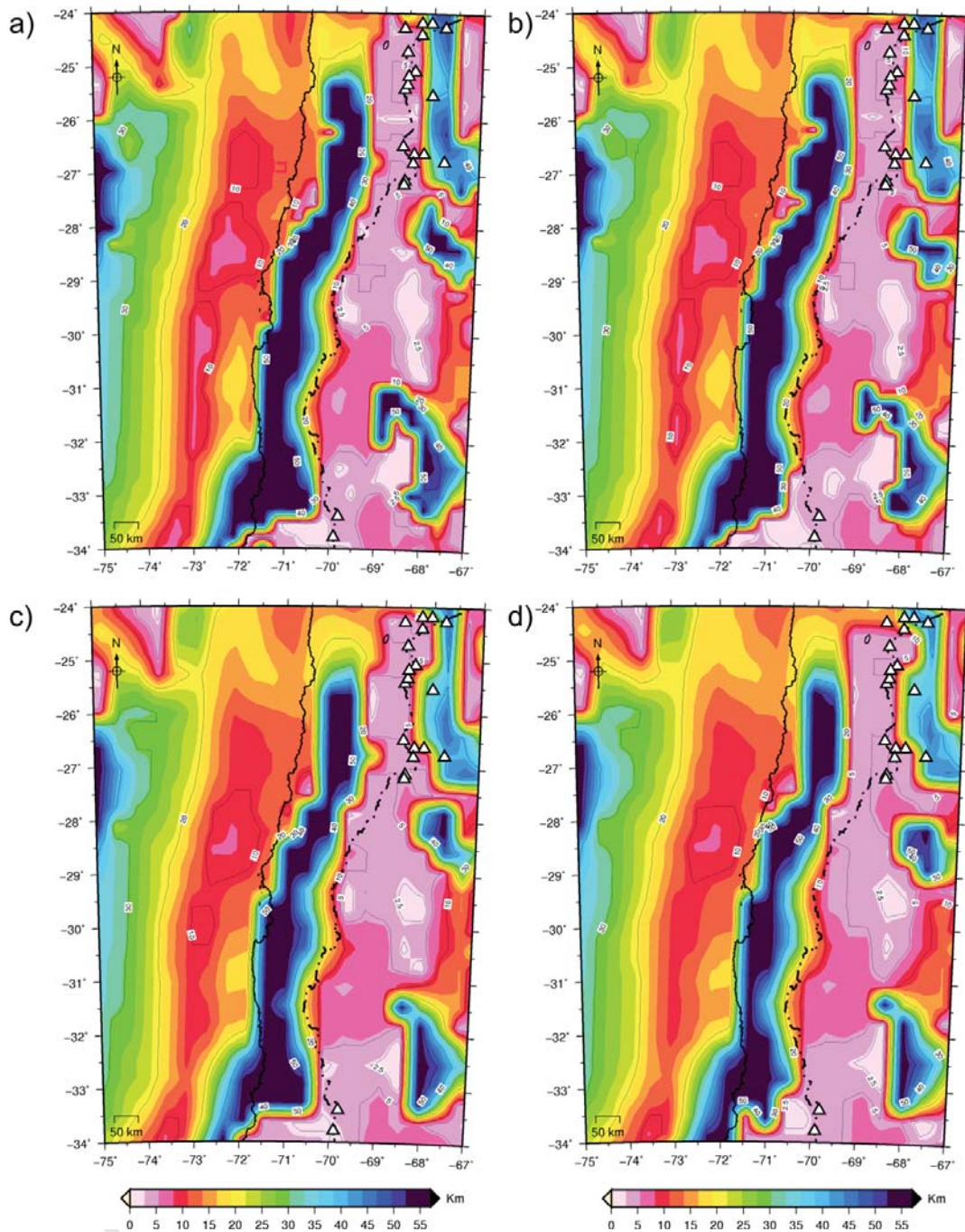
1437 **Figure C3.** Histogram of the residual Moho between the crust–mantle interface
 1438 obtained by gravity inversion and the crust–mantle interface from the flexural model.
 1439 Globally, more than 85% of the error is less than 4 km. **a)** Window size 60x60km (error
 1440 $\approx 89\% < 4$ km). **b)** Window size 70x70km. (error $\approx 88\% < 4$ km) **c)** Window size

79

1441 $80 \times 80 \text{ km}$ (error $\approx 87\% < 4 \text{ km}$) **d)** Window size $90 \times 90 \text{ km}$ (error $\approx 85\% < 4 \text{ km}$). The
1442 residual obtained for the smaller window size of $60 \times 60 \text{ km}$ shows a slight better fitting,
1443 and increasing error as window size increases. We selected the window size of $80 \times 80 \text{ km}$
1444 as is in the order of magnitude of the half wavelength of the resolution of GOCE model.

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1447 **Figure C4.** Elastic thicknesses obtained from GOCE data for different window size of
 1448 calculation. **a)** Window size 60x60km **b)** Window size 70x70km. **c)** Window size
 1449 80x80km **d)** Window size 90x90km.

1450

81

1451 From the reduced Bouguer anomaly (Figure 6a) we estimated the gravimetric crust-
1452 mantle discontinuity (Figure 6b) by gravity inversion. This method uses an iterative
1453 algorithm that alternates downward continuation with direct forward modeling
1454 (Braitenberg et al. 1999) and is somewhat analogous to the Oldenburg-Parker inversion
1455 approach (Oldenburg 1974; see Braitenberg et al. 2007, for a detailed explanation).

1456 The obtained crustal load (obtained from topo/bathymetry data and the density model)
1457 and the Moho undulations (obtained by inversion of the reduced Bouguer anomaly)
1458 were used for the inverse flexure calculation. The flexural rigidity was inverted in order
1459 to match the known loads with the known crustal thickness model (i.e. to model the
1460 gravity Moho in terms of an isostatic model). The elastic thickness was allowed to vary
1461 in the range of $1 < T_e < 55$ km and was iteratively estimated over moving windows of 60
1462 x 60 km, 70 x 70 km, 80 x 80 km and 90 x 90 km, size (Fig. C4). The model
1463 parameters are given in Table 1, where the adopted densities are standard values already
1464 used in the region under study by other authors (Gimenez et al. 2000, 2009; Introcaso et
1465 al. 2000; Miranda and Robles, 2002).

1466 The residual between the Moho from gravity inversion and the flexure Moho is the
1467 residual Moho (Fig. C2). The Moho undulations obtained from gravity inversion agree
1468 with the CMI undulations expected for the flexural model, about 85 percent within 4 km
1469 of difference (Fig. C3). Positive values of the residual Moho indicate that the gravity
1470 Moho is shallower than the flexure Moho. The correction of the gravity and load effect
1471 of the sediments, allows gravity Moho to comply with load changes: once the negative
1472 effect on the Bouguer anomaly from the sediments is removed, the Moho from gravity
1473 inversion will be shallower and will follow the flexure Moho (Alvarez et al. 2015).
1474 Negative values of the residual Moho in the Main Andes indicate that the gravity Moho

1475 is deeper than the flexure Moho. The flexural model used in this work is a
1476 simplification and would be influenced by the stress of the down-going plate. Thus, the
1477 Te solutions along the active subduction margin could possibly be distorted (as
1478 explained by Braitenberg et al. 2007).

1479

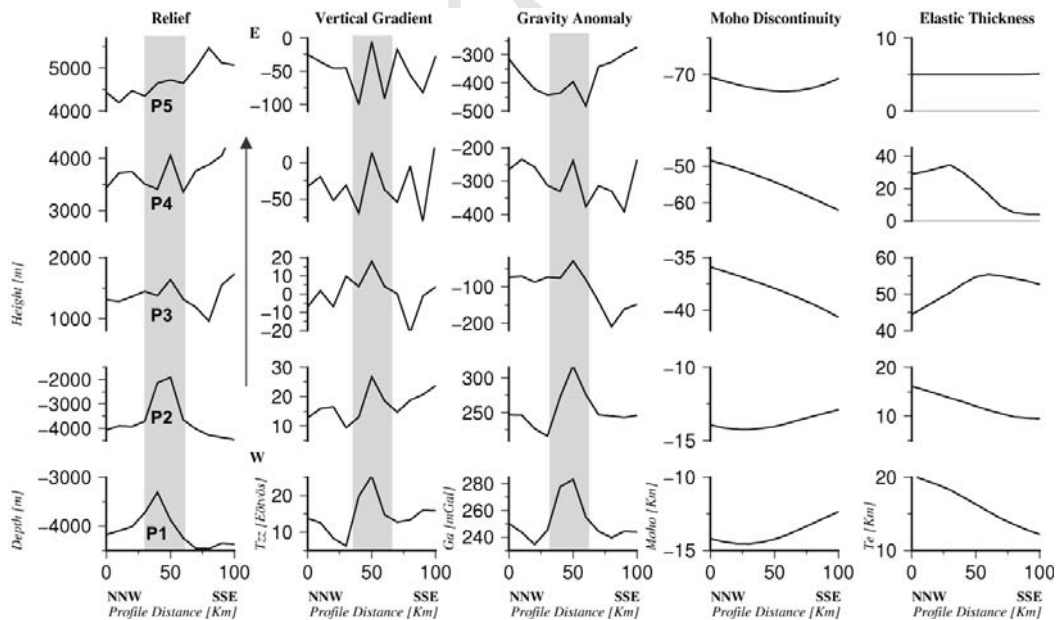
1480 **APPENDIX D: TERRAIN FABRIC ANALYSIS**

1481 The terrain fabric or grain (Pike et al. 1989) measures a point property of a digital
1482 elevation model (DEM) or any other field (in this case is applied to the Tzz),
1483 constituting an expression of the tendency to form linear features (Guth, 1999). Guth
1484 (1999) explained that eigenvector analysis reliably extracts the terrain fabric, both in
1485 strength and orientation. This analysis extracts eigenvectors and eigenvalues (S) from a
1486 3 by 3 matrix of the sum of the cross products of the directional cosines of the surface
1487 perpendiculars at each point in the DEM (Guth, 1999). The dominant orientation is
1488 represented by the vector V1 that together with the vector V2 forms the main plane of
1489 the fabric, while V3 is normal to this plane. The eigenvalues S1, S2 and S3 express their
1490 normalized values. The $\ln(S2/S3)$ measure the fabric or grain, orientations of S2 and S3
1491 define the dominant grain of the topography, and the ratio S2/S3 determines the strength
1492 of the grain. The eigenvalue method was discussed by Woodcock (1977) for the
1493 representation of fabric shapes in structural geology, paleomagnetism, sedimentology
1494 and glaciology. This method is used to quantify Chapman's (1952) technique in
1495 Microdem software (Guth, 1995, 2007).

1496

1497 **APPENDIX E: Gravity derived profiles**

1498 Five profiles (Fig. E1) were traced perpendicularly to the Copiapó ridge track (for
 1499 location see Fig. 5b) for the T_{zz} and Ga from EGM2008 and for Moho and Te from
 1500 GOCE. While profiles P1 and P2 were traced over the Copiapó ridge in an offshore
 1501 position, profiles P3, P4 and P5 were traced cutting across the Ojos del Salado-San
 1502 Buenaventura volcanic lineament (Fig. 5b). The bathymetric expression of the Copiapó
 1503 ridge (P1, P2) is expressed by high values of T_{zz} and Ga . Similarly, inland, the
 1504 topographic highs along the ridge path (P3, P4, P5) are also reflected by high values of
 1505 T_{zz} and Ga . Even though the profile P5 does not show a significant topographic
 1506 expression, a relatively high T_{zz} value was also obtained. The Ga (P5) shows a relative
 1507 high too, although it is masked by the negative effect of the Andean root that is
 1508 expressed by the low Moho values.



1509

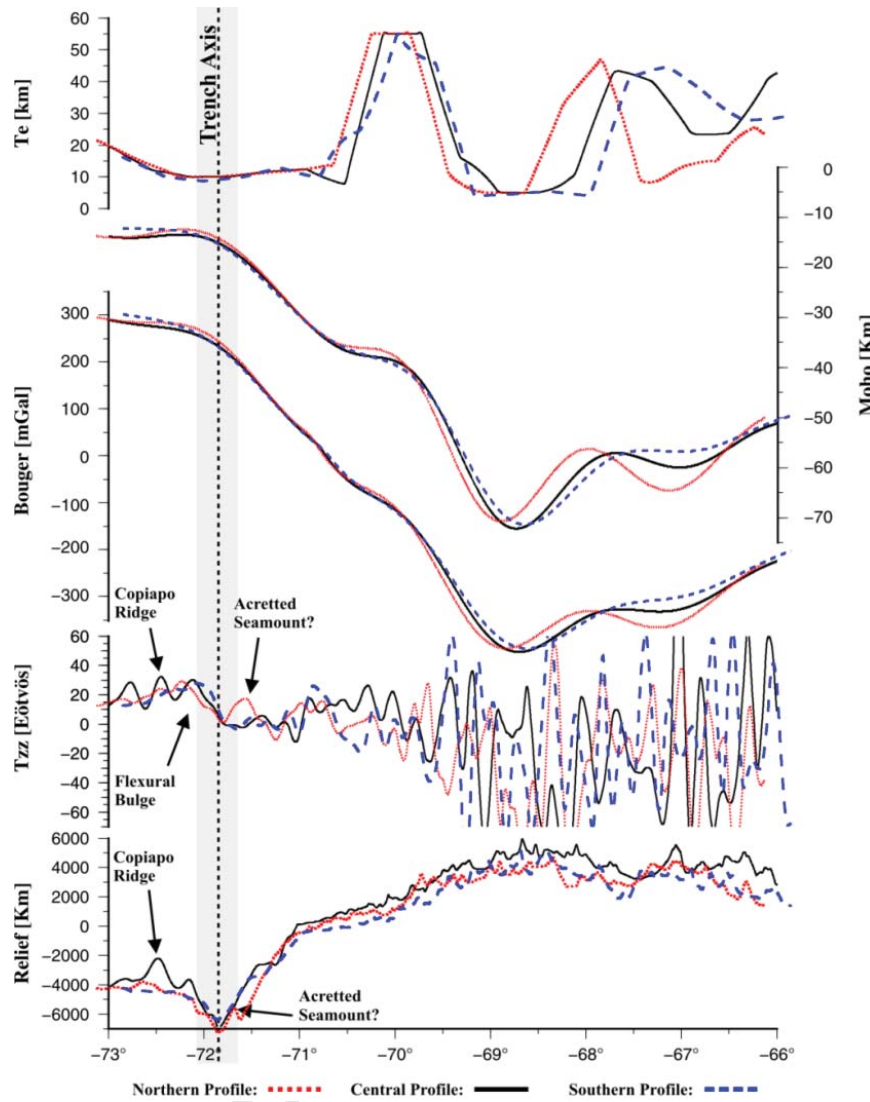
1510 **Figure E1.** Longitudinal profiles (C1, Northern profile, C2, Central profile, C3,
 1511 Southern profile) at the path of the Copiapó ridge offshore and beneath the South

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1512 *America Plate, comparing topography, T_{zz} , Bouguer anomaly, Moho and T_e (see*
1513 *Figure 5b for location).*

1514

1515 Additionally, three profiles were traced in the N-NE direction along and parallel to the
1516 ridge track (for location see Fig. 5b). Profile C1 (Fig. E2, red dotted line) was traced
1517 parallel and to the north of the ridge path; Profile C2 (Fig. E2, black solid line) was
1518 traced over the Copiapó ridge and through the Ojos del Salado-San Buenaventura
1519 volcanic lineament (Fig. 5b); and the southern profile C3 was traced parallel and to the
1520 south of C2 (Fig. E2, blue dashed line). In a general analysis, profile C2 clearly shows
1521 the roughed and prominent shape of the Copiapó ridge, while the other profiles present
1522 a flat bottom bathymetry. Additionally, the T_{zz} signal reflects the ridge track (C2), and
1523 the positive effect of the flexural bulge (outer rise) (C1, C2, C3). Inland, the profile C2
1524 shows in general a higher mean value of topography along its track.



1525

1526 **Figure E2.** 2D profiles across the Copiapó ridge path comparing topography, T_{zz} ,
 1527 Bouguer anomaly and T_e (see Figure 5b for location).

1528 **Table A1:** Euler poles used in the reconstruction.

Moving/Fix Plate	Lon	Lat	Age(My)	Angle	Reference
Naz./South-Am.	-98°	-54.9°	8	5.328	<i>DeMets et al. 2010</i>
Naz./Pac.	-89.75°	60.11°	10.9	-14.88	<i>Muller et al. 2008</i>

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	-91.50°	64.5°	20.1	-30.70	<i>Muller et al. 2008</i>
Far./Pac.	-92.61°	73.53°	23.0	-31.08	<i>Muller et al. 2008</i>
	-110.70°	76.10°	33.1	-45.27	<i>Tebens and Cande, 1997</i>
	-122.97°	82.19°	40.1	-52.24	<i>Muller et al. 1997</i>
	-160.16°	85.24°	47.9	-62.78	<i>Muller et al. 1997</i>

1529

Table 1. Euler poles used in the reconstruction.

1530

1531 **Highlights**

1532 Vertical gravity gradient show a deflection of the main Andean structures trend

1533 Relationship between the Copiapo ridge and deformation in the upper plate

1534 The Copiapo ridge controls the northern edge of the Chilean flat subduction zone

1535 Leading to shallower subduction at the northern region of the Chilean flat-slab

1536 Chilean-Pampean flat slab could be related to two simultaneous ridge collision

1537

Moving/Fix Plate	Lon	Lat	Age(Ma)	Angle	Reference
Naz./South-Am.					
	-98°	-54.9°	8	5.328	DeMets et al., 2010
Naz./Pac.					
	-89.75°	60.11°	10.9	-14.88	Muller et al., 2008

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Far./Pac.	-91.50°	64.5°	20.1	-30.70	Muller et al., 2008
	-92.61°	73.53°	23.0	-31.08	Muller et al., 2008
	-110.70°	76.10°	33.1	-45.27	Tebens and Cande, 1997
	-122.97°	82.19°	40.1	-52.24	Muller et al., 1997
	-160.16°	85.24°	47.9	-62.78	Muller et al., 1997

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Table 1. Euler poles used in the reconstruction.

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<i>Average difference</i>	<i>0.0772 mgal</i>
<i>Standard Deviation</i>	<i>12.341 mgal</i>
<i>Maximal value of difference</i>	<i>62.021 mgal</i>

1543

Table B1: Statistical parametrs for the difference.

1544

1545

Masses above sea level	ρ_s	2,670 kg/m ³
Upper crustal density	ρ_{uc}	2,700 kg/m ³
Lower crustal density	ρ_{lc}	2,900 kg/m ³
Upper mantle density	ρ_m	3,300 kg/m ³

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Young modulus	E	10^{11} N/m ²
Poisson ratio	Σ	0.25

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Table C1. Parameters used in the flexural modeling.

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